



Australia's National Science Agency

# Simulation of the ASTRI demonstration particle receiver during on-sun testing

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1. CSIRO Energy, Newcastle, NSW, Australia

Asia Pacific Solar Research Conference 2023

Wednesday 6<sup>th</sup> of December 2023

RMIT, Melbourne

Land of the Wurundjeri people of the Kulin Nations



ASTRI

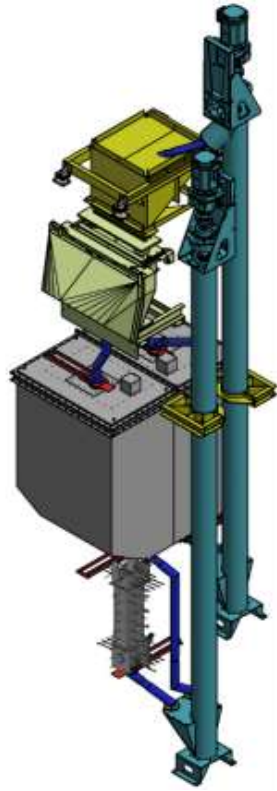
Australian Solar Thermal  
Research Institute



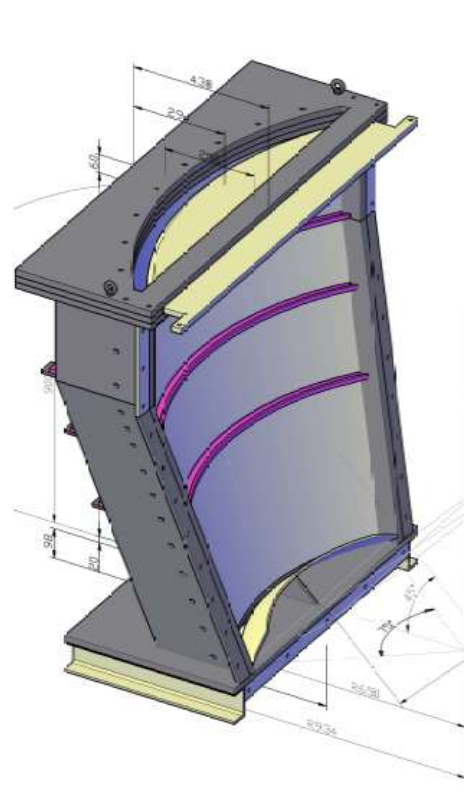
# ASTRI demonstration particle receiver



(a) Solar Field 2 with particle receiver system at CSIRO Newcastle



(b) CAD model of particle receiver system with cooler



(c) CAD model of particle receiver (casing removed)

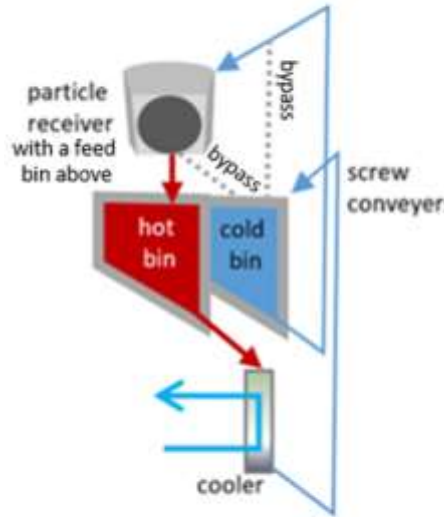


(d) Particle falling test

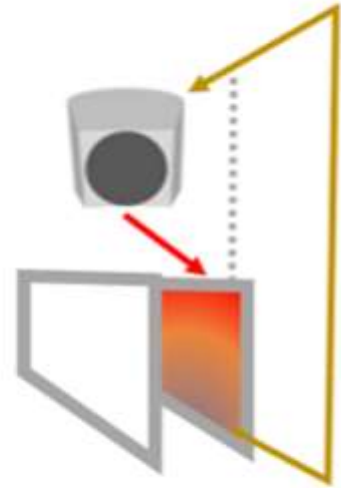
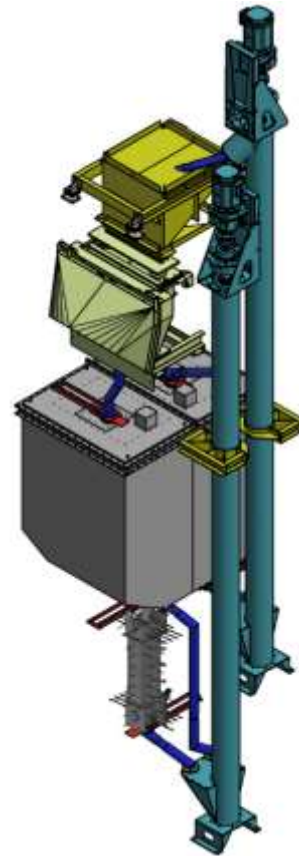


# ASTRI demonstration particle receiver

- A 500kW packed-bed heat exchanger for cooling the particles for two-bin operation (left) has been installed
- Experiments considered in the present work (May 2022 to March 2023) used single-bin operation (right) where particles are recirculated without cooling



(a) Two-bin operation with particle cooler

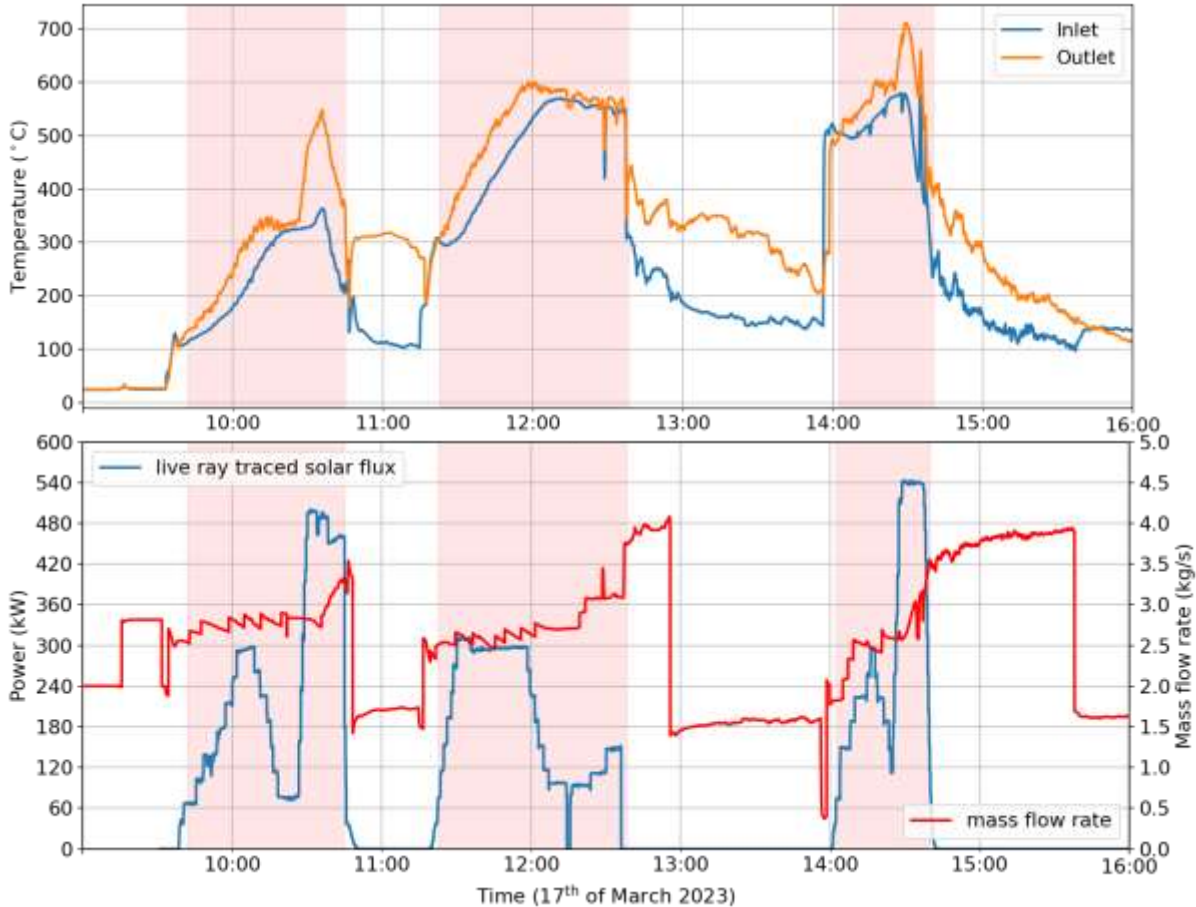


(b) Single-bin operation (present work)



# Experiment data

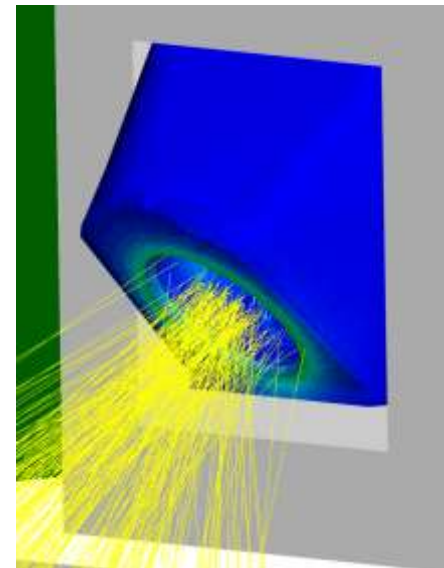
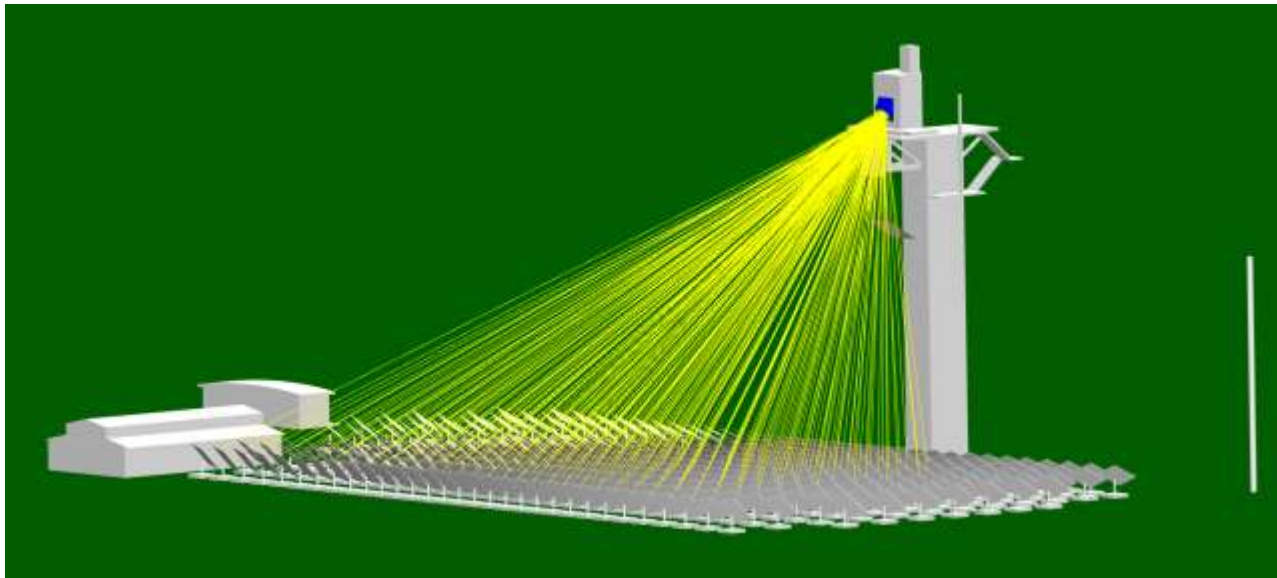
- Key parameters required for computational rebuilding are recorded during experiments:
  - Particle temperature at inlet and outlet
  - Particle mass flow rate
  - DNI
  - Helio­stat field utilisation
- Typical day involves operating the receiver for multiple ~1hr runs with breaks in between
- Outlet temperatures reached up to 700°C with  $\Delta T \approx 100^\circ\text{C}$
- Goal of present work is to rebuild experiments with Heliosim and compare with the measured data





# Physical modelling

- Heliostat field optics and receiver heat transfer model created using CSIRO's Heliosim software [1,2]



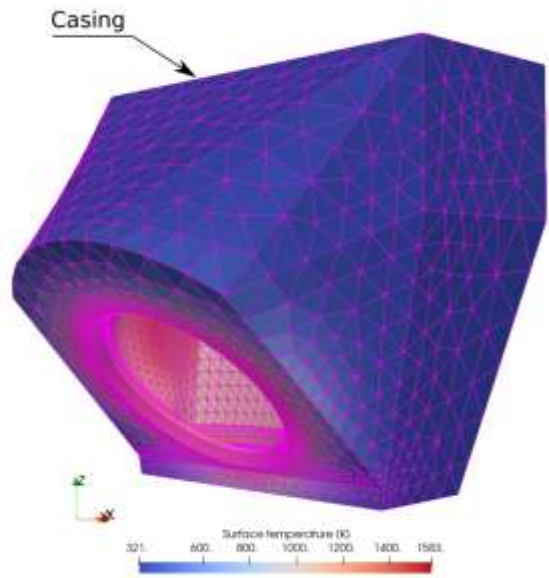
1. D. Potter et al., AIP Conference Proceedings, Nov. 2018, vol. 2033, no. 1, p. 210011, doi: 10.1063/1.5067213.

2. D. Potter, "The Heliosim software family: modelling, design, and control tools for CST systems," in Proceedings of the Asia Pacific Solar Research Conference, 2022.

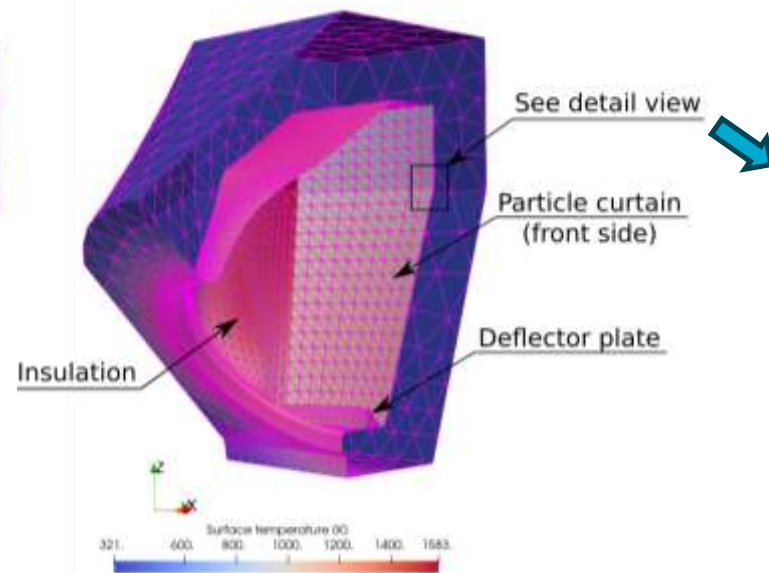


# Receiver modelling

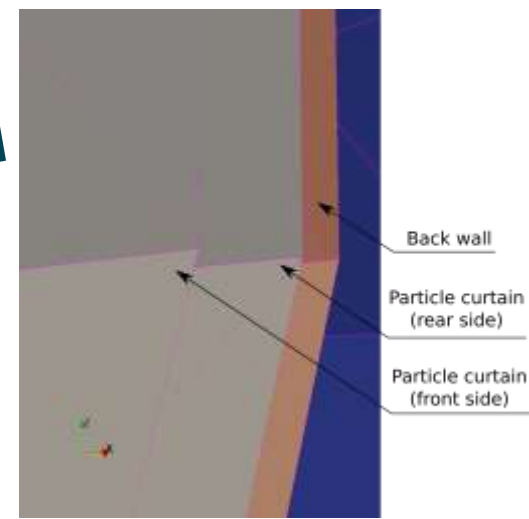
- Surface mesh -based model of receiver with front and rear sides of particle curtain
- Steady state energy balance applied to each mesh facet to solve for temperature



(a) Receiver surface mesh  
( $12 \times 10^3$  facets)



(b) Receiver surface mesh  
clipped in symmetry plane



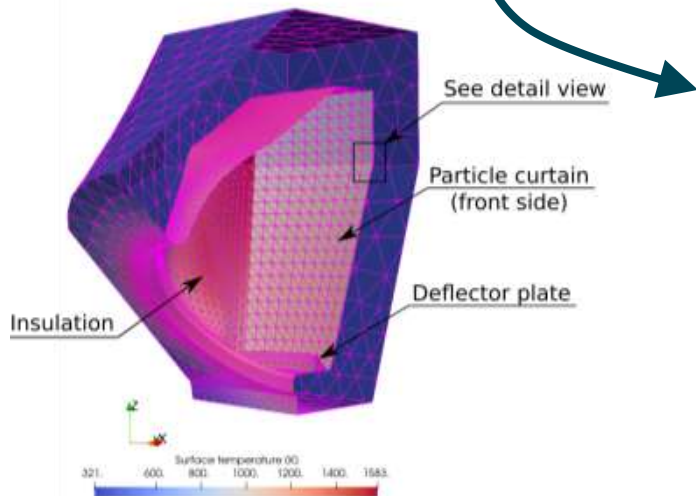
(c) Particle curtain and back  
wall detail



# Receiver modelling

## Receiver heat transfer mechanisms

- Solar and thermal radiation heat transfer
- Convective heat loss due to ambient air flow
- Conduction through walls
- Transport of thermal energy by particle advection



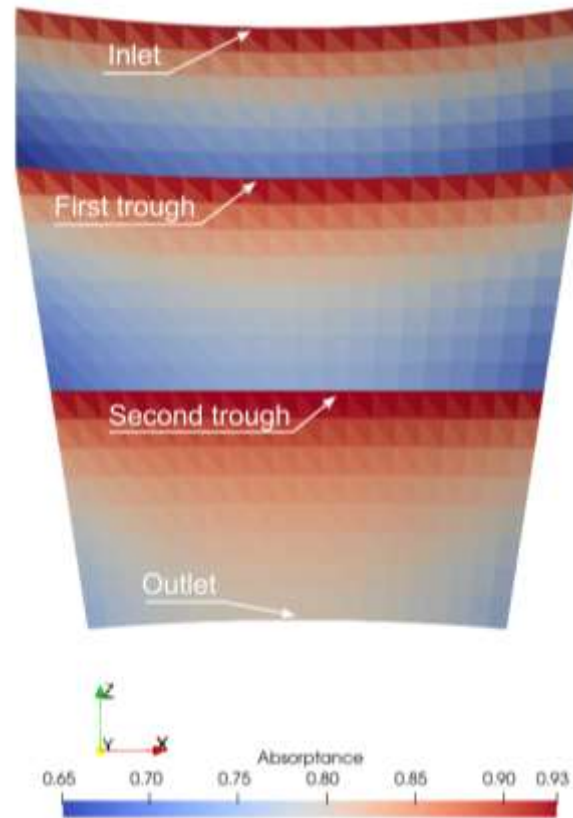
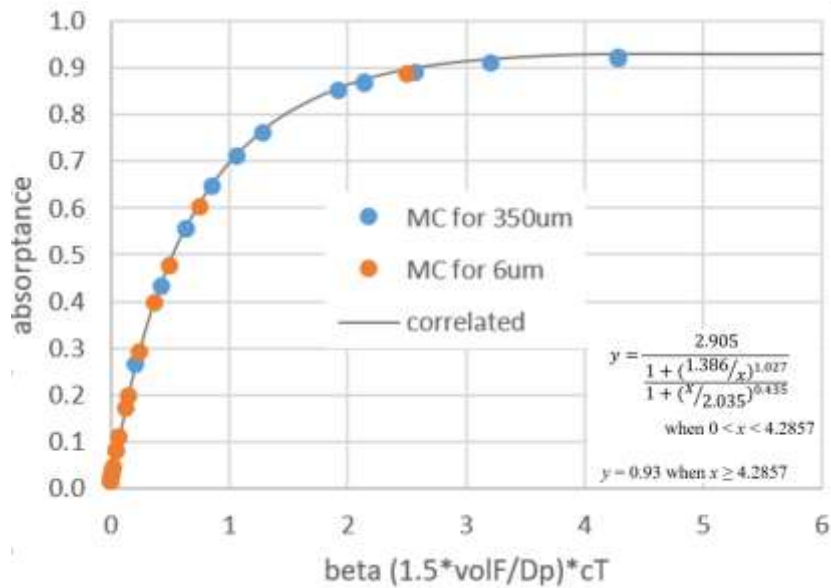
## Particle curtain model

- Divided into multiple vertical “channels” with equal mass flow rate
  - Mixing between adjacent channels is assumed to be negligible
  - Drag not considered
- Constant temperature assumed through the curtain thickness
- Optical properties are calculated using correlations



# Curtain optical properties

- Correlations fitted to detailed Monte Carlo ray tracing simulations [1] for various curtain thicknesses, particle diameters and particle volume fractions



[1] A. Kumar et al., *J. Sol. Energy Eng. Trans. ASME*, vol. 140, no. 6, 2018, doi: 10.1115/1.4040290.

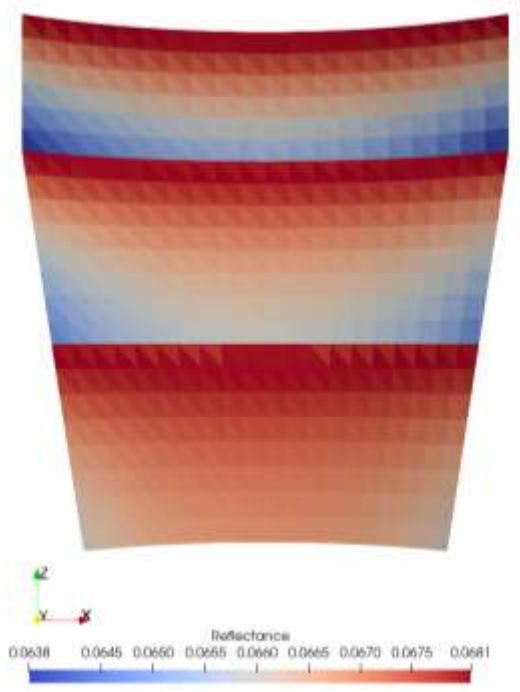
One-way solar absorptance at design point (2.73kg/s)



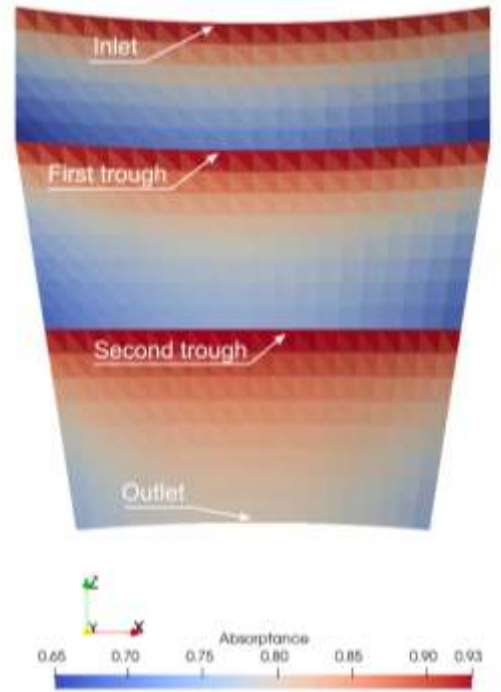


# Curtain optical properties

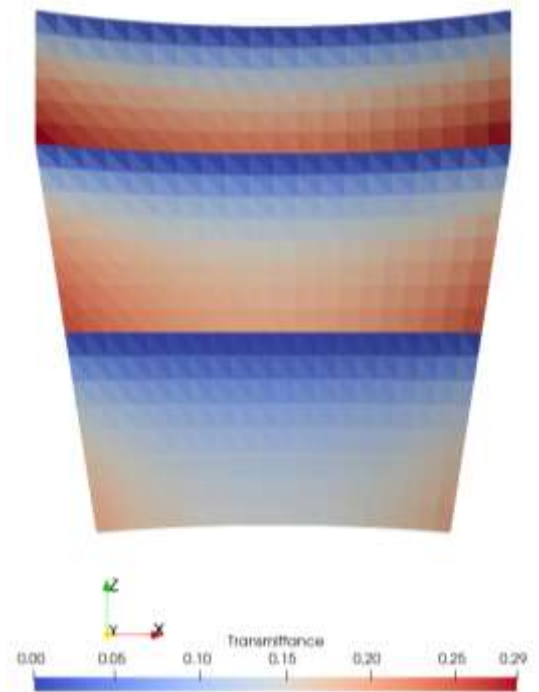
- One-way optical properties at design point (2.73kg/s):



(a) Diffuse reflectance  
(average of 0.0668)



(b) Absorptance  
(average of 0.8039)

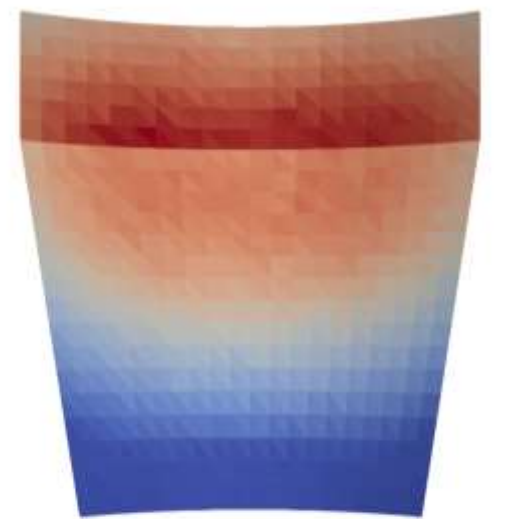


(c) Transmittance  
(average of 0.1292)

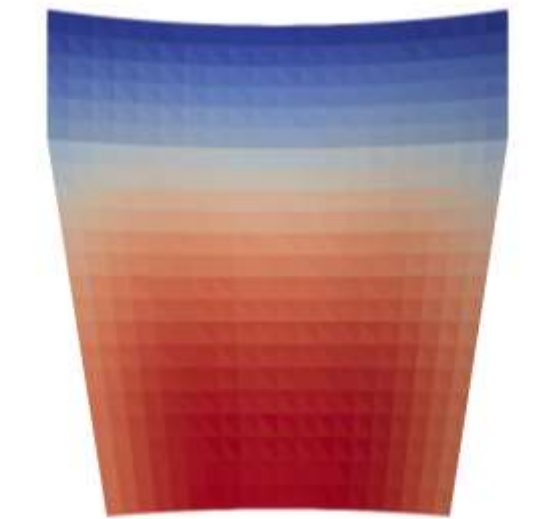


# Curtain optical properties

- Particle curtain and back wall temperature distributions at design point conditions:



(a) Particle curtain incident solar irradiance



(b) Particle curtain temperature

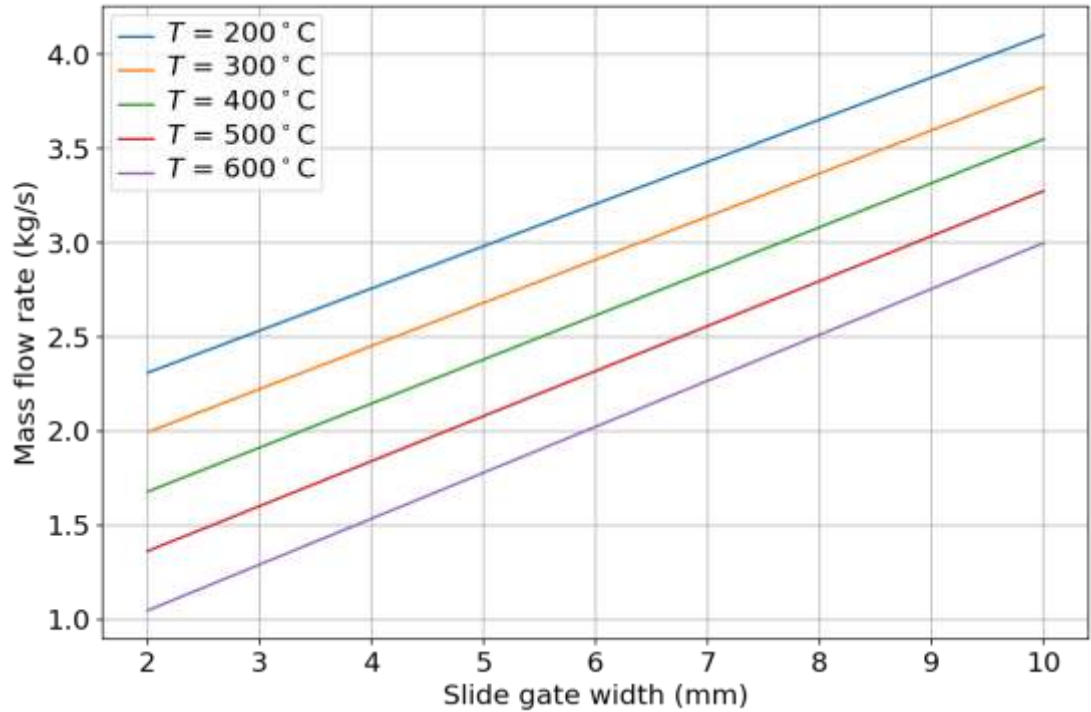


(c) Back wall temperature



# Mass flow rate correlation

- Designed method for mass flow rate determination was measuring change in feed hopper mass over a fixed time interval using a load cell
- Thermal expansion issues made this method unreliable at high temperatures ( $\geq 400^{\circ}\text{C}$ )
- After this issue was fixed in 2023, a correlation was made to calculate mass flow rate as a function of slide gate width and particle inlet temperature [1]

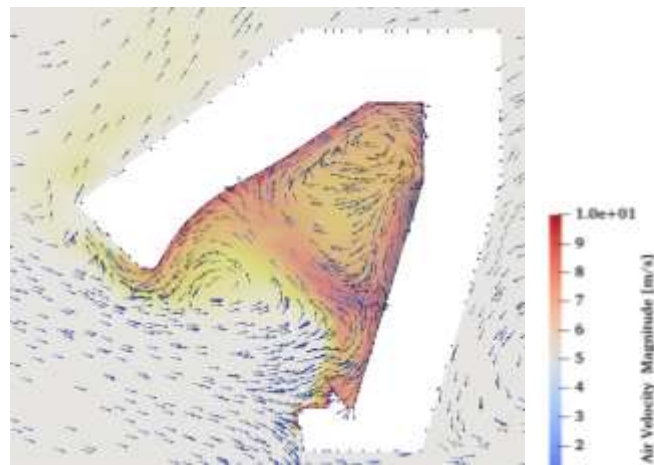
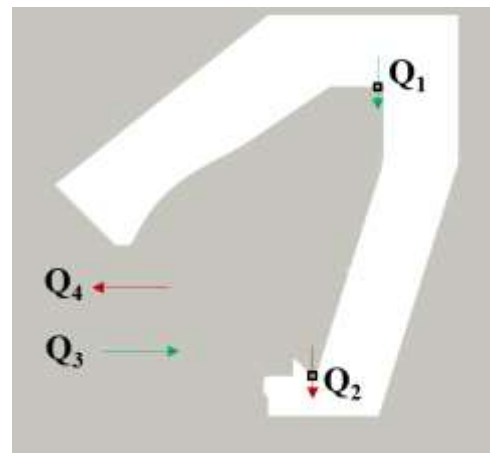
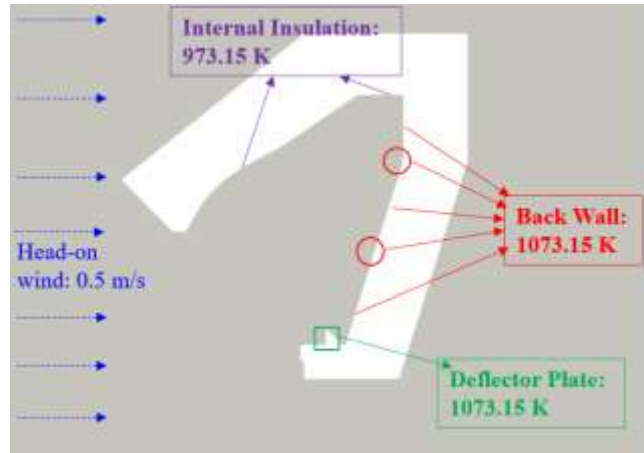


[1] G. R. Drewer, J. Kim, and D. Potter, "The Importance of Managing the Performance of Particle Lift and Flow Control Systems in Research CST Plants to Facilitate Commercialization," presented at SolarPACES 2023.

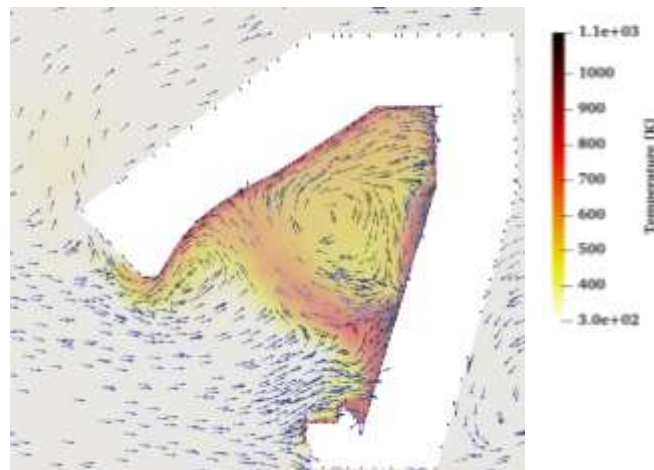


# Convective heat loss

- ‘Coarse-grained’ CFD-DPM simulations performed using OpenFOAM [1]
- Fixed wall and particle temperatures
- Particle flow rate was varied from 5 million particles/s (0.4 kg/s) to 30 million particles/s (2.4 kg/s)
- Net convective heat loss computed as  $Q_4 + Q_2 - Q_1 - Q_3$



(a) 10 million particles/s (0.81 kg/s)



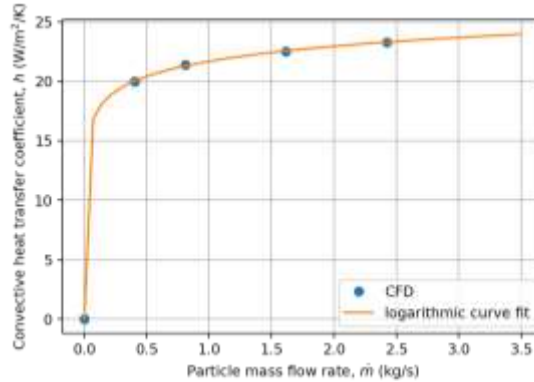
(b) 20 million particles/s (1.62 kg/s)

# Convective heat loss coefficients

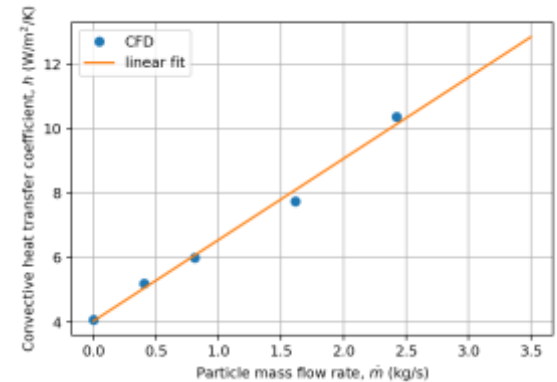
- Convective heat loss coefficients for each receiver surface computed from CFD-DPM heat flux results:

$$h = \frac{q}{A(T - T_a)}$$

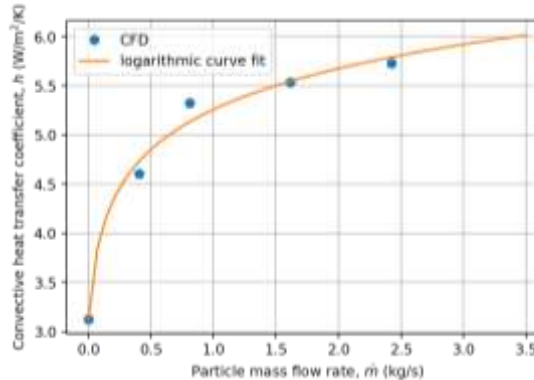
- Correlations for convective heat loss coefficients as a function of mass flow rate determined by fitting to CFD-DPM results using logarithmic expressions



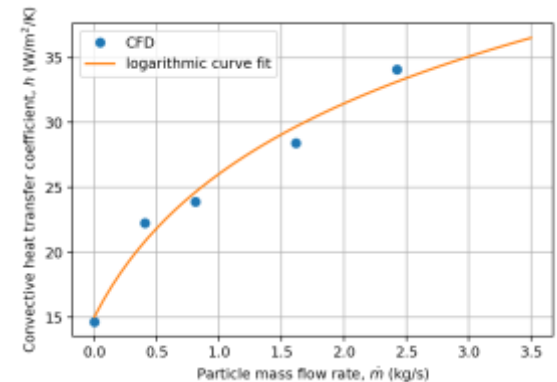
(a) Particle curtain (single-sided area)



(b) Back wall



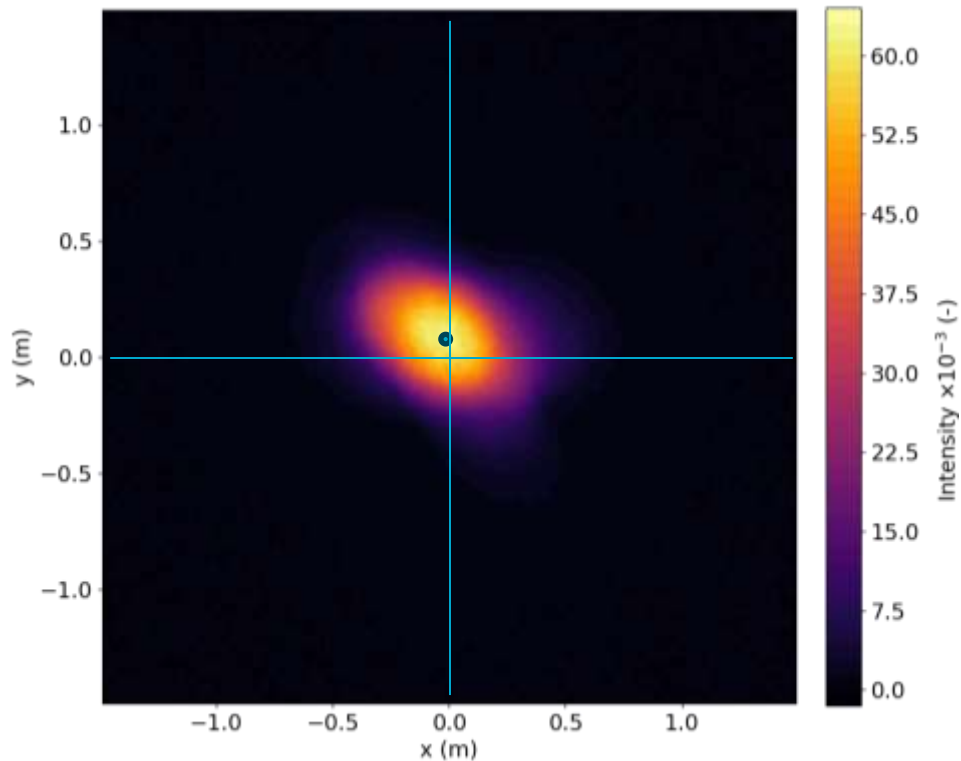
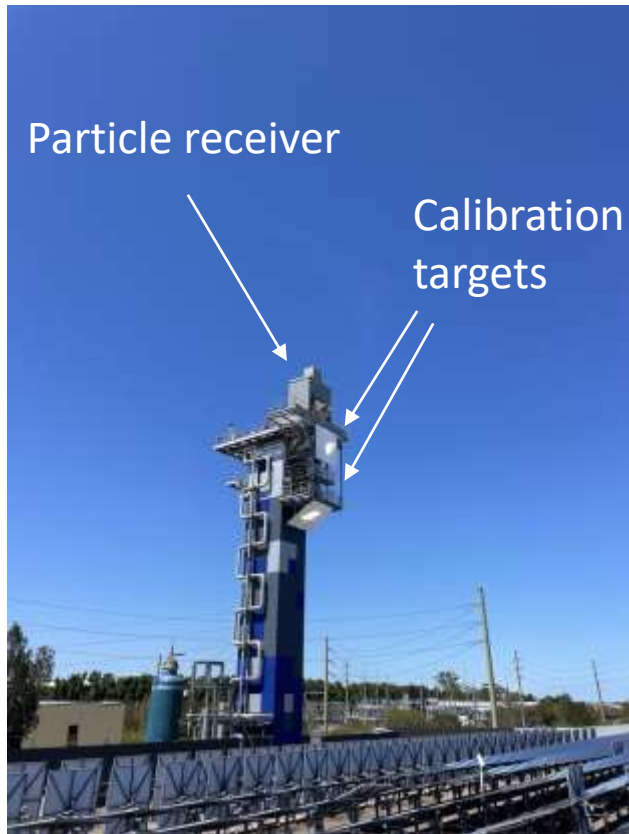
(c) Insulation



(d) Deflector plate



# Tracking error measurements



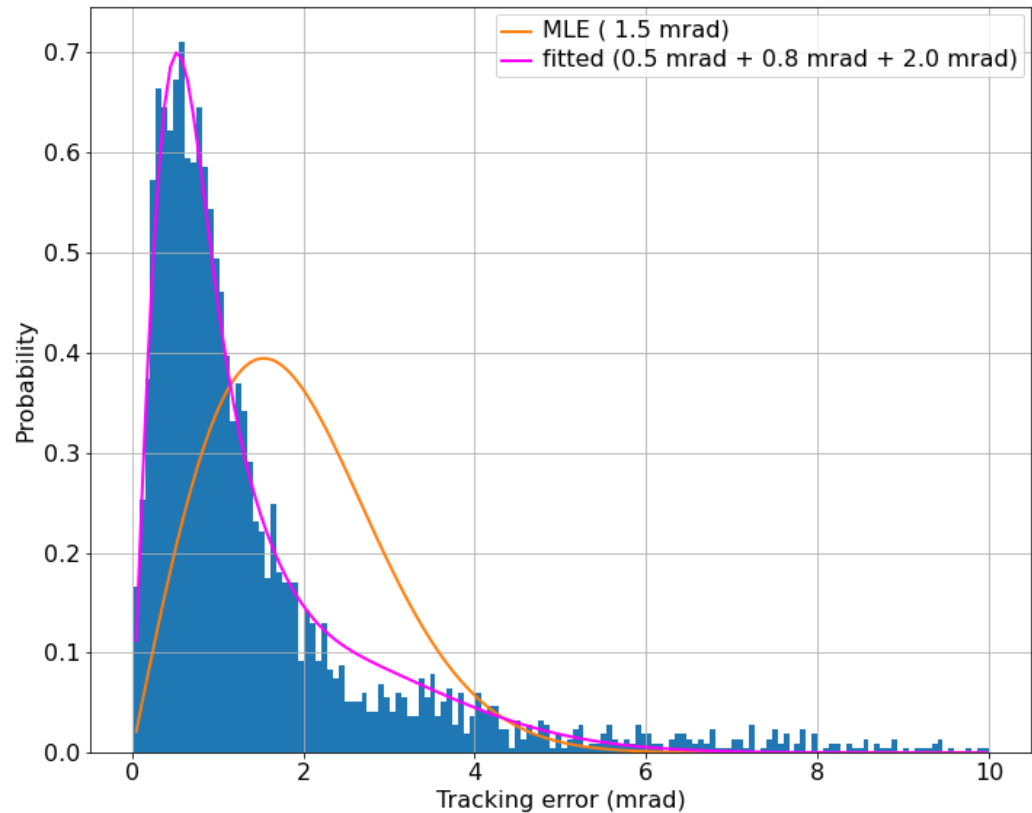
Typical heliostat image on target just prior to calibration procedure





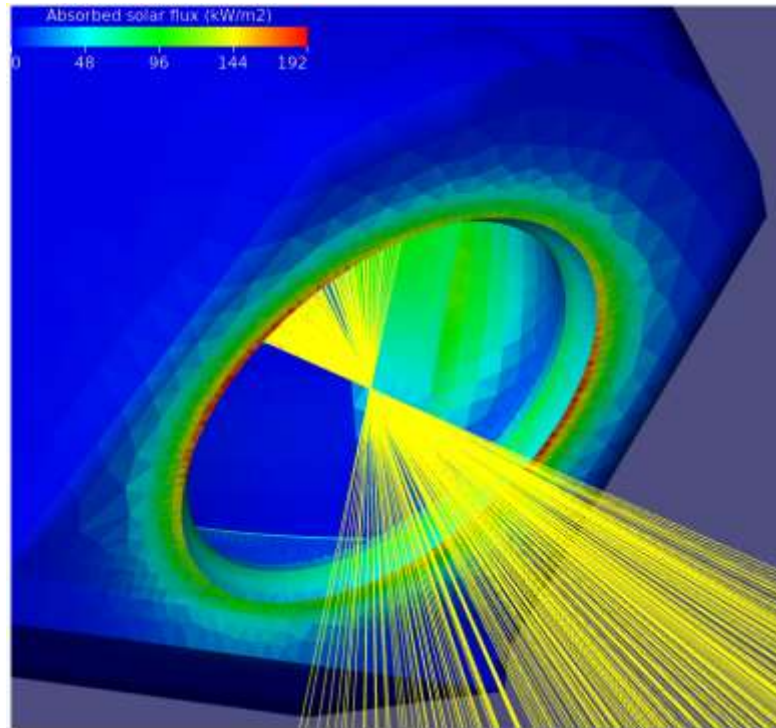
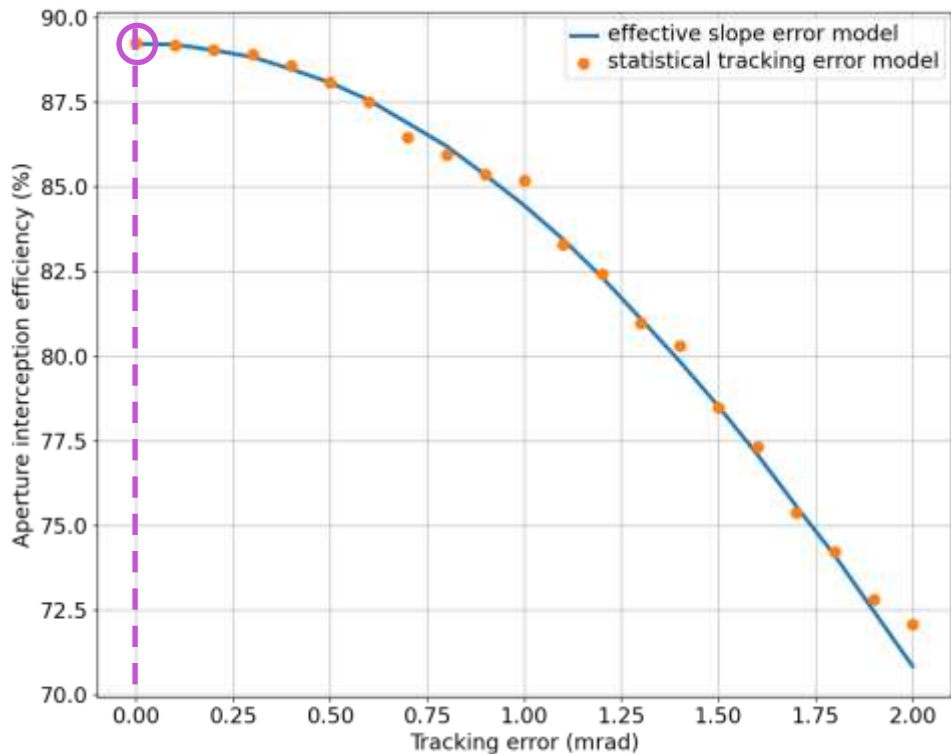
# Tracking error measurements

- 3314 calibration images acquired between November 2021 and May 2023
- Tracking errors greater than 10mrad discarded (56 cases)
- Probability histogram is well described by a combination of multiple (3) evenly weighted Rayleigh distributions
- Multiple Rayleigh distribution model has yet to implemented in Heliosim, therefore tracking error is varied between 0.5 and 1.5mrad to gauge sensitivity





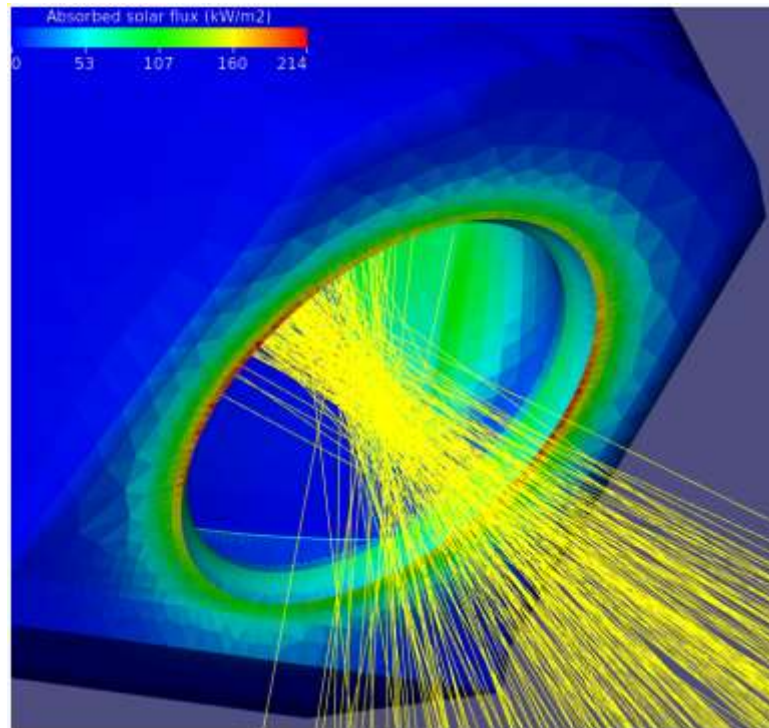
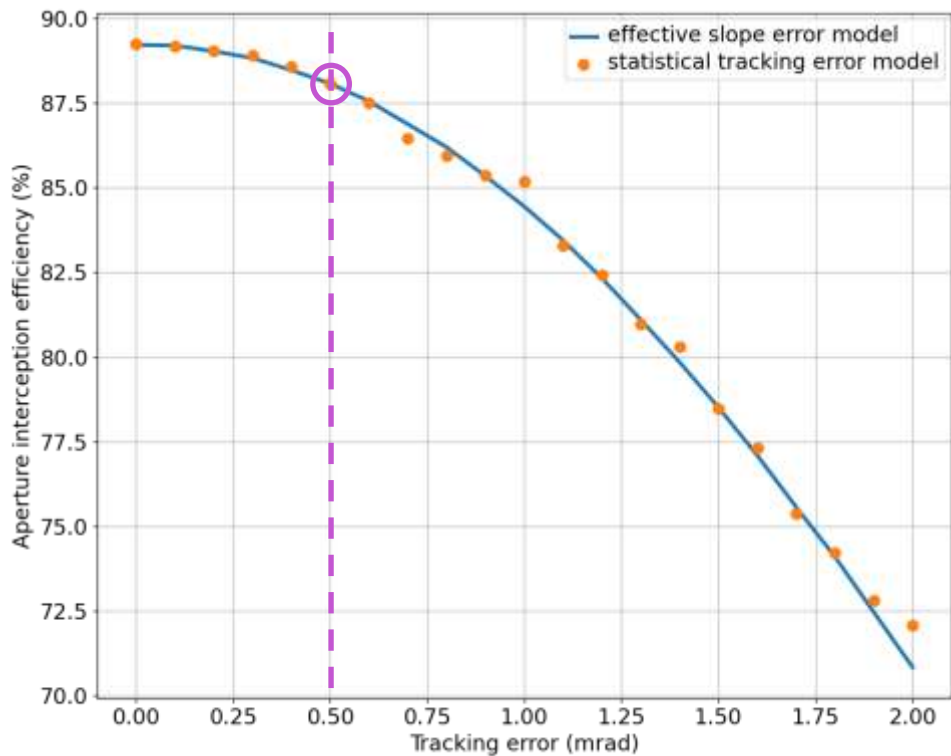
# Tracking error modelling



0 mrad



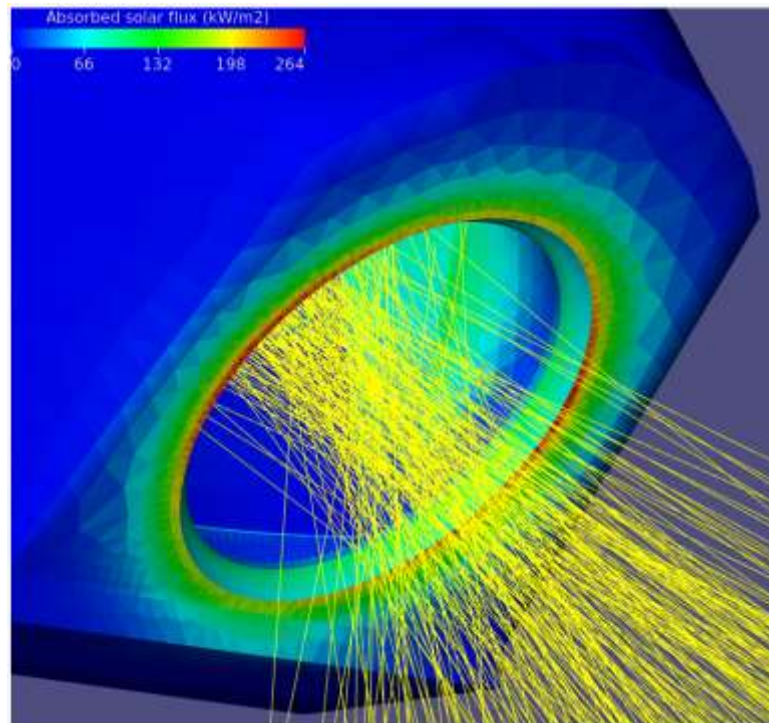
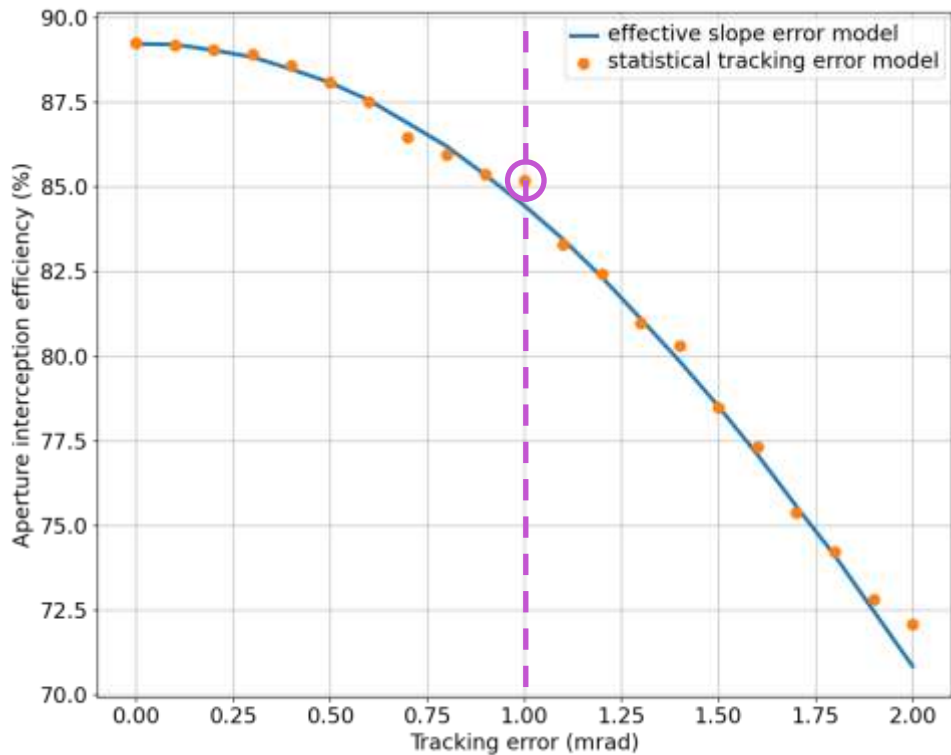
# Tracking error modelling



0.5 mrad



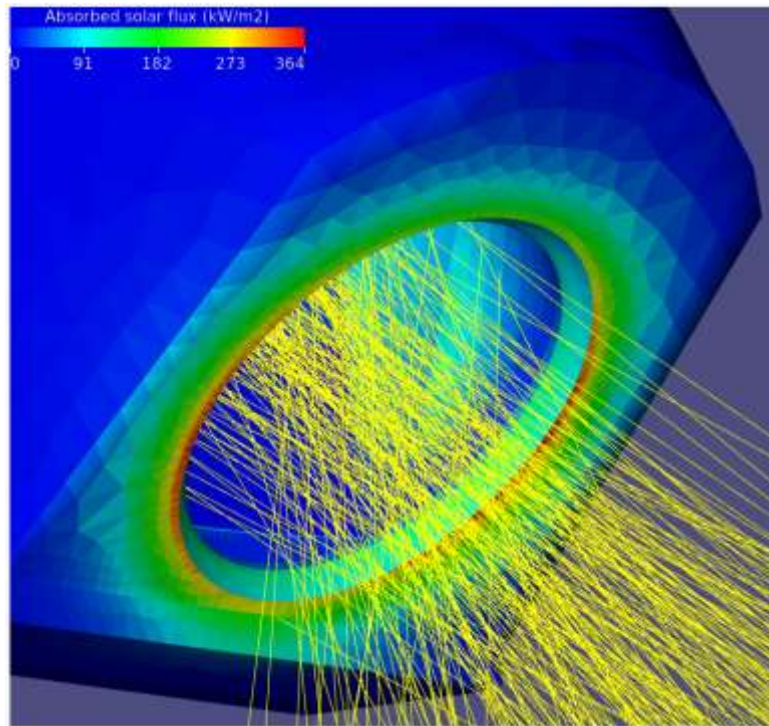
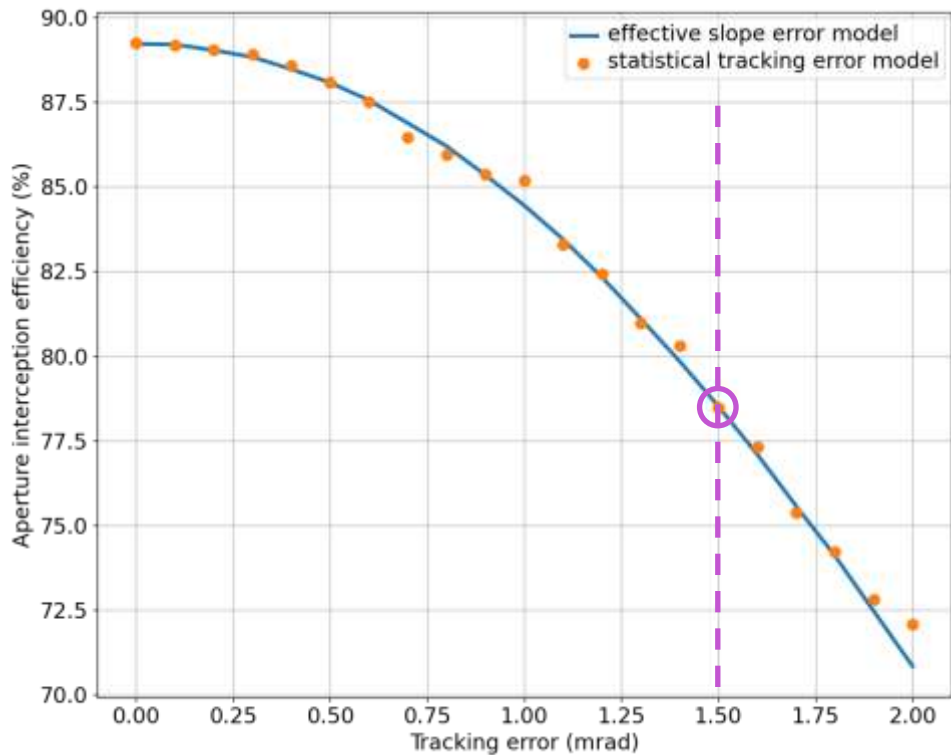
# Tracking error modelling



1.0 mrad



# Tracking error modelling

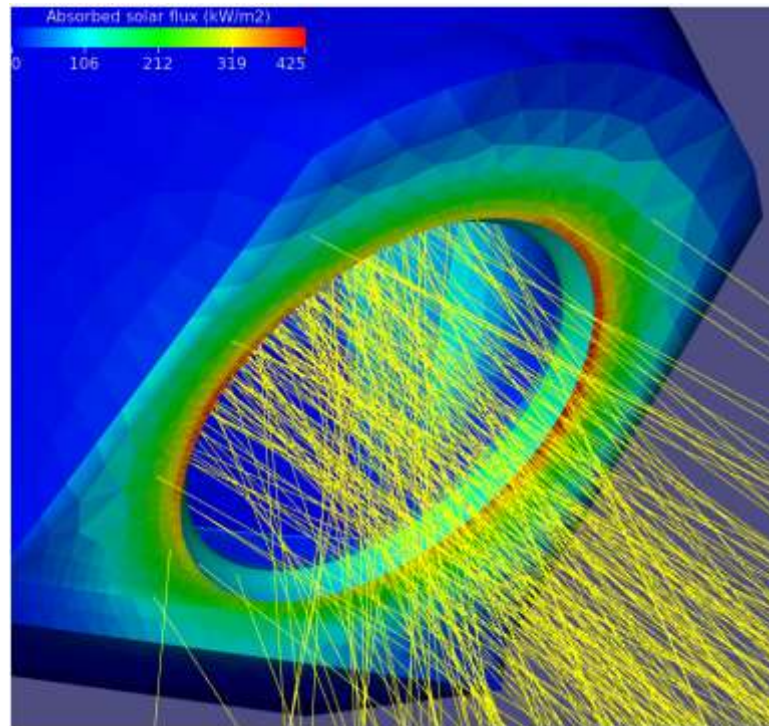
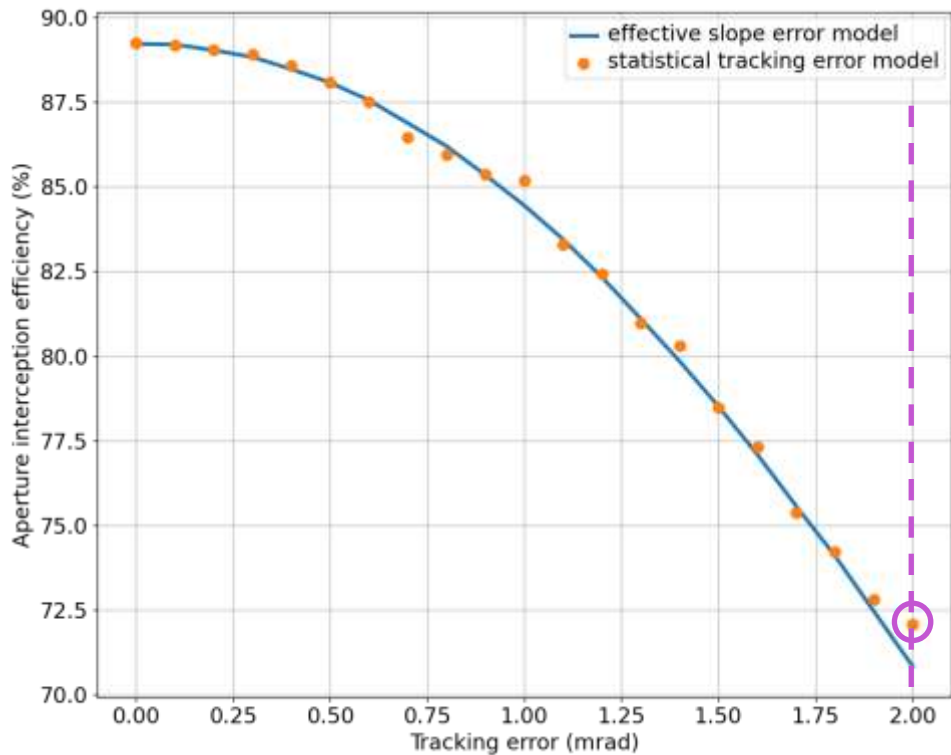


1.5 mrad





# Tracking error modelling



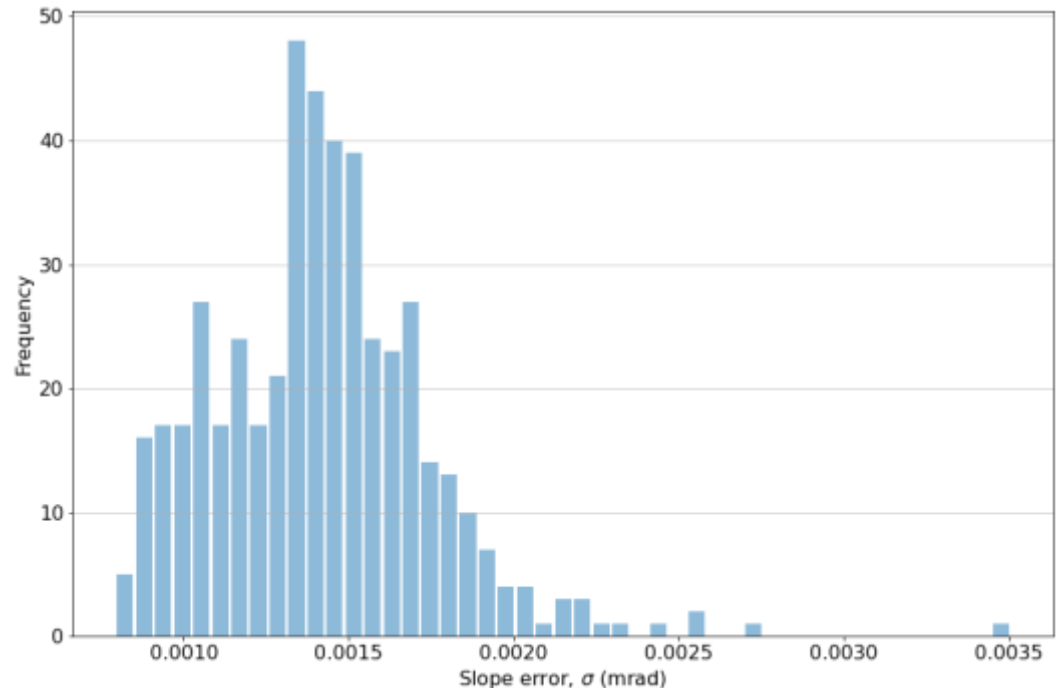
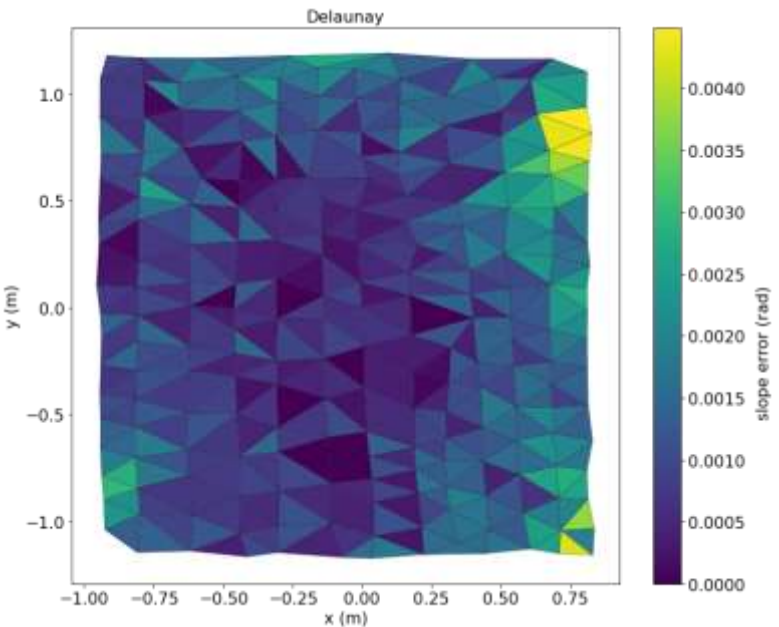
2.0 mrad





# Heliostat facet modelling

- 90% specular reflectance assumed
- Slope error of 1.4mrad used based on analysis of surface metrology data for Solar Field 2 heliostats

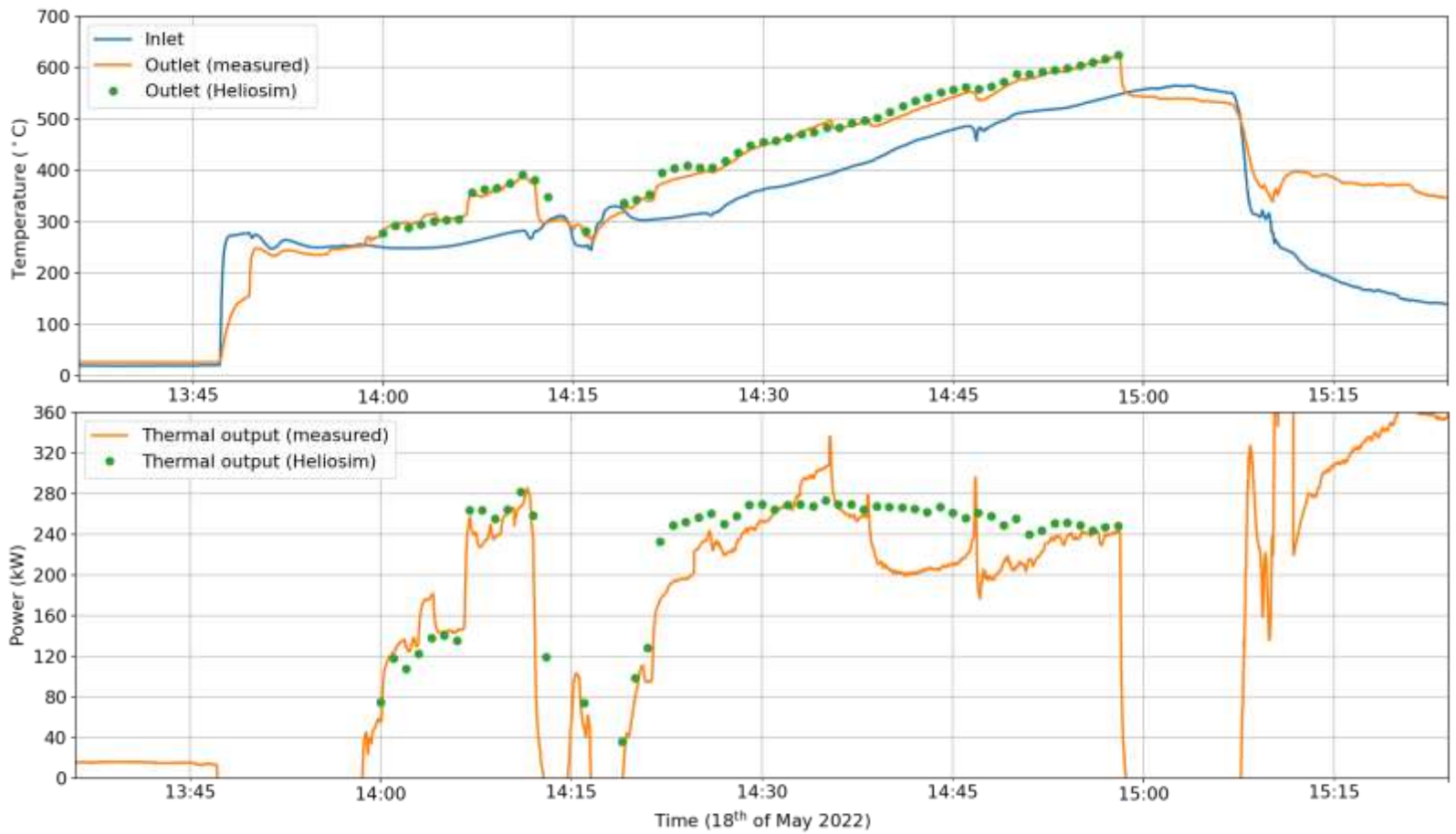


(a) Example of surface mesh constructed from measured facet coordinates

(b) Histogram of average facet slope error for Solar Field 2 heliostats

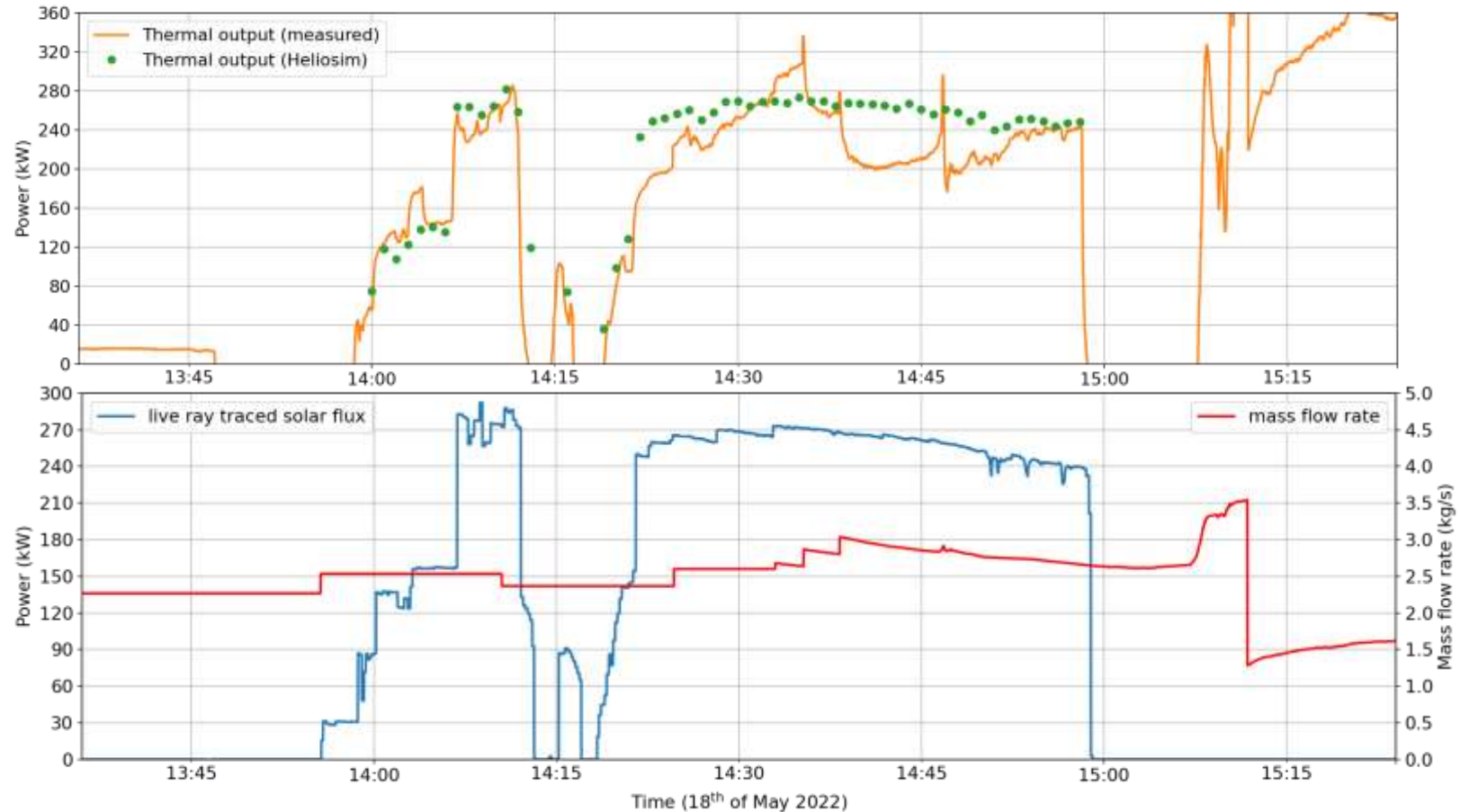


# Results – 18<sup>th</sup> of May 2022



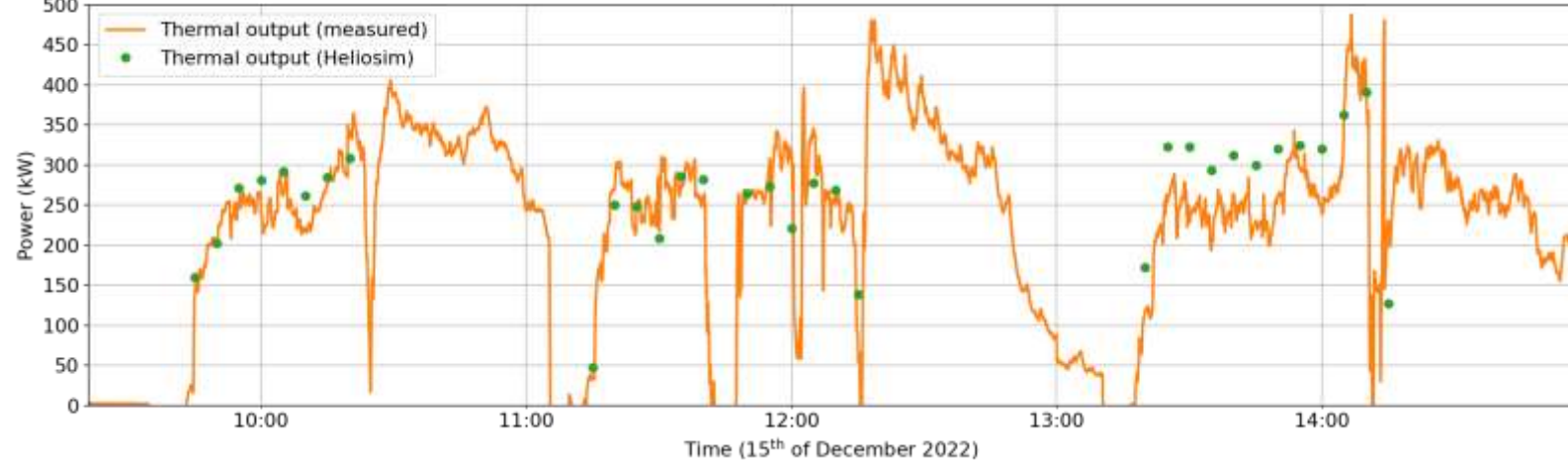
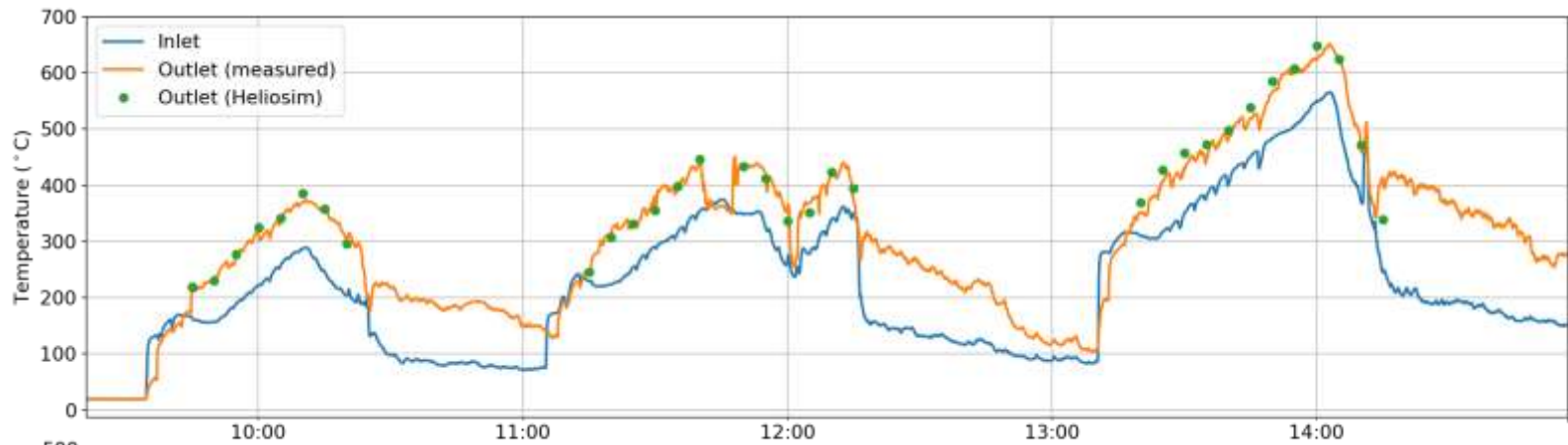


# Results – 18<sup>th</sup> of May 2022



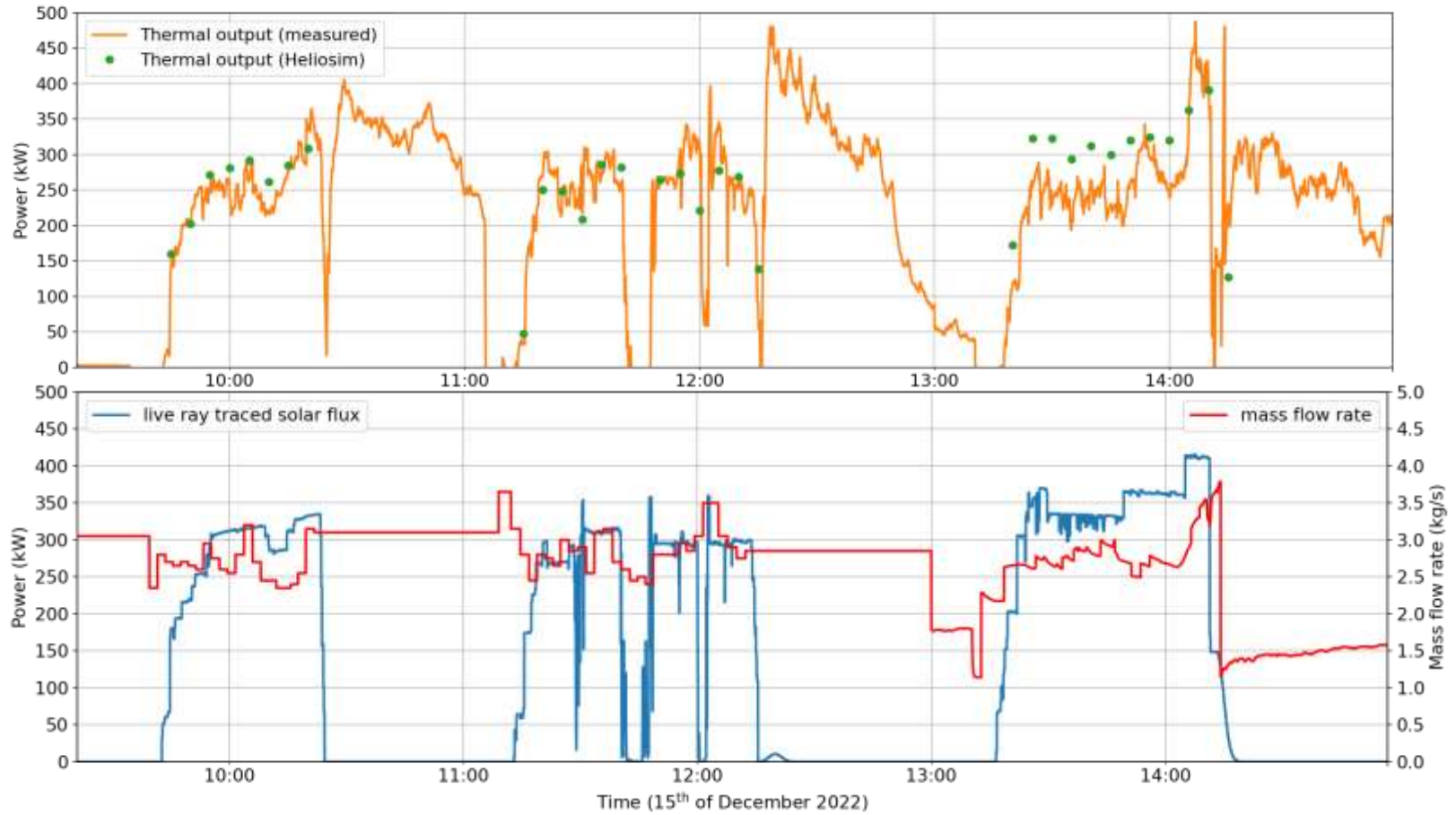


# Results – 15<sup>th</sup> of December 2022



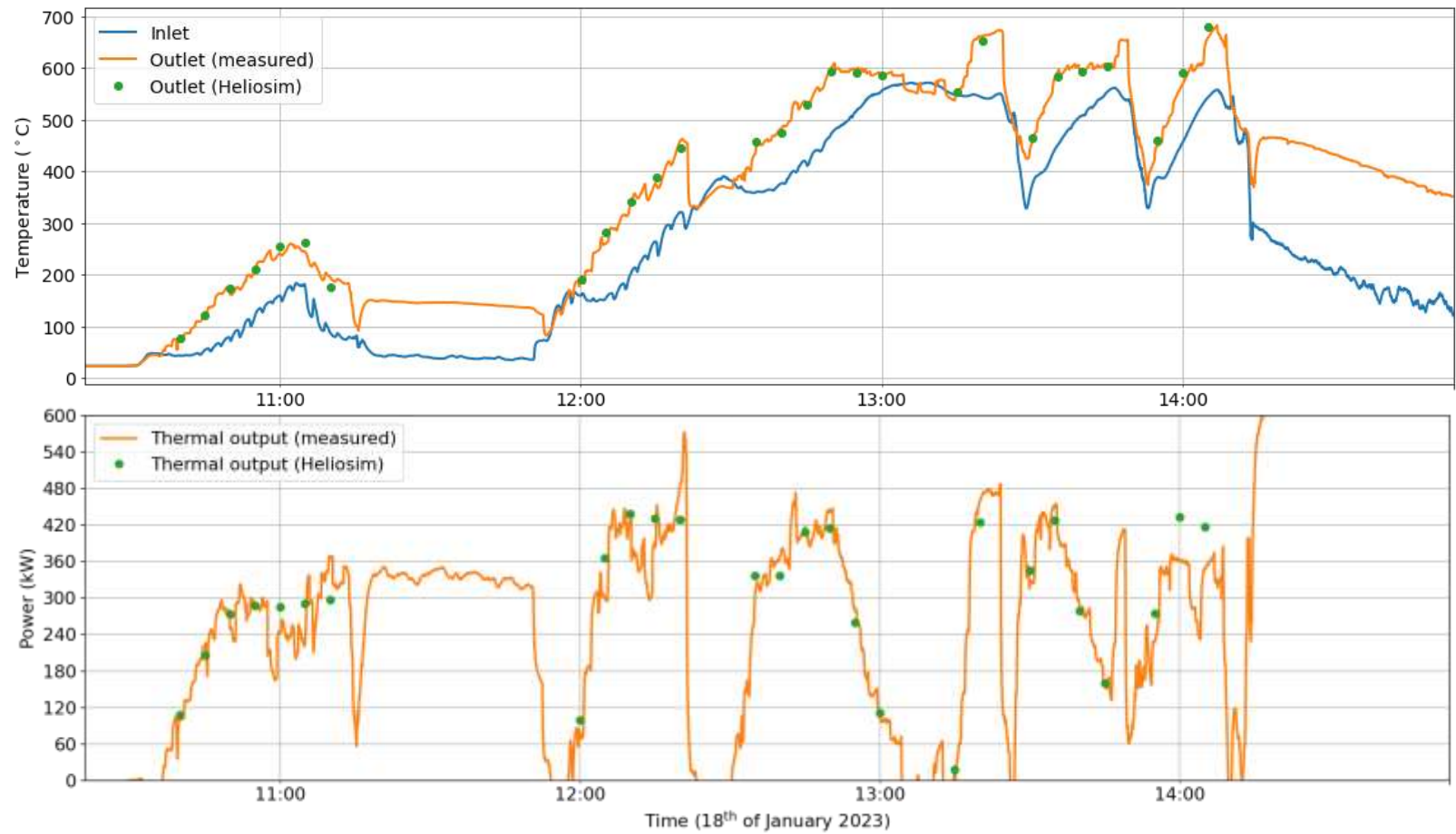


# Results – 15<sup>th</sup> of December 2022





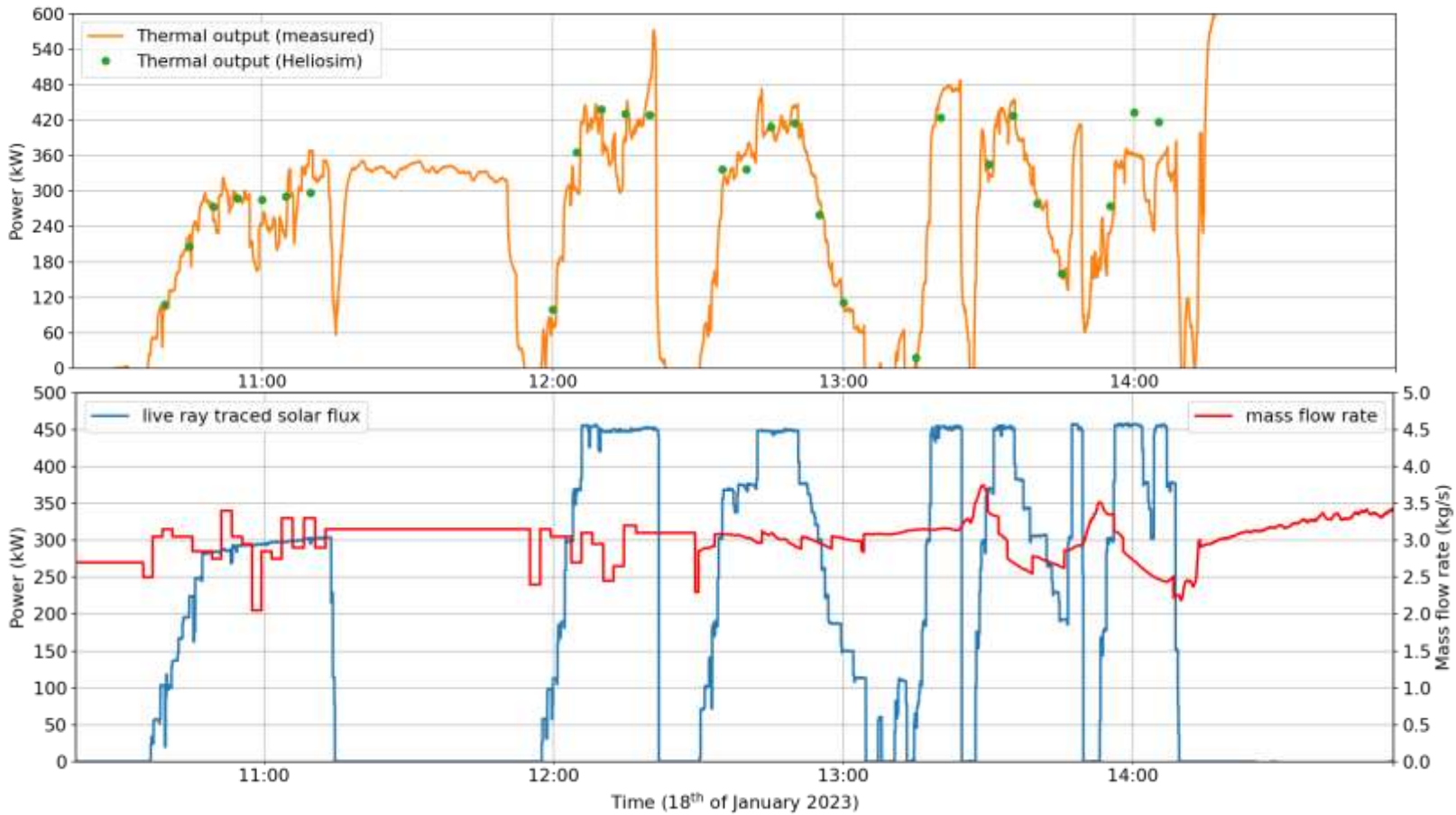
# Results – 18<sup>th</sup> of January 2023





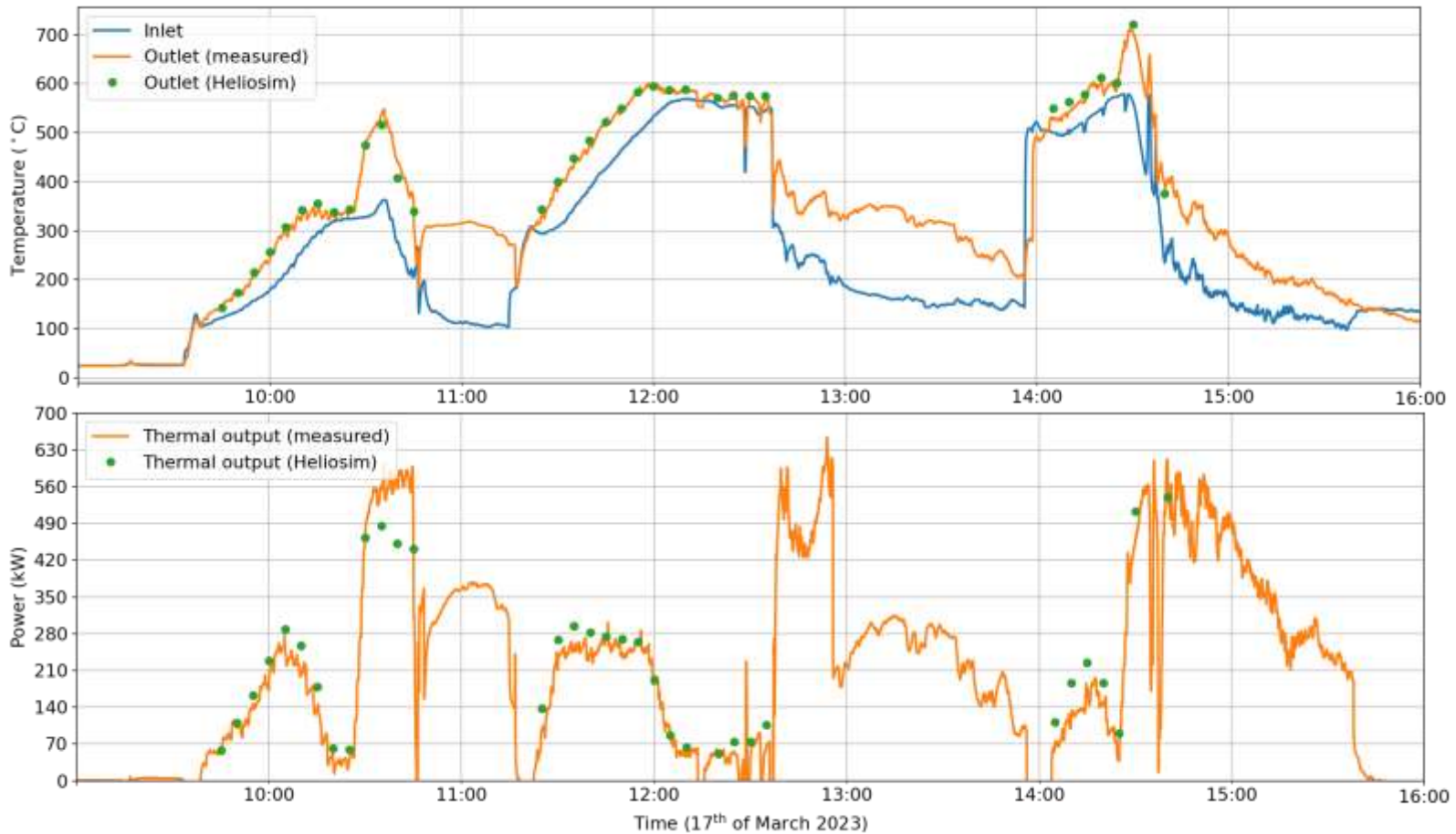


# Results – 18<sup>th</sup> of January 2023



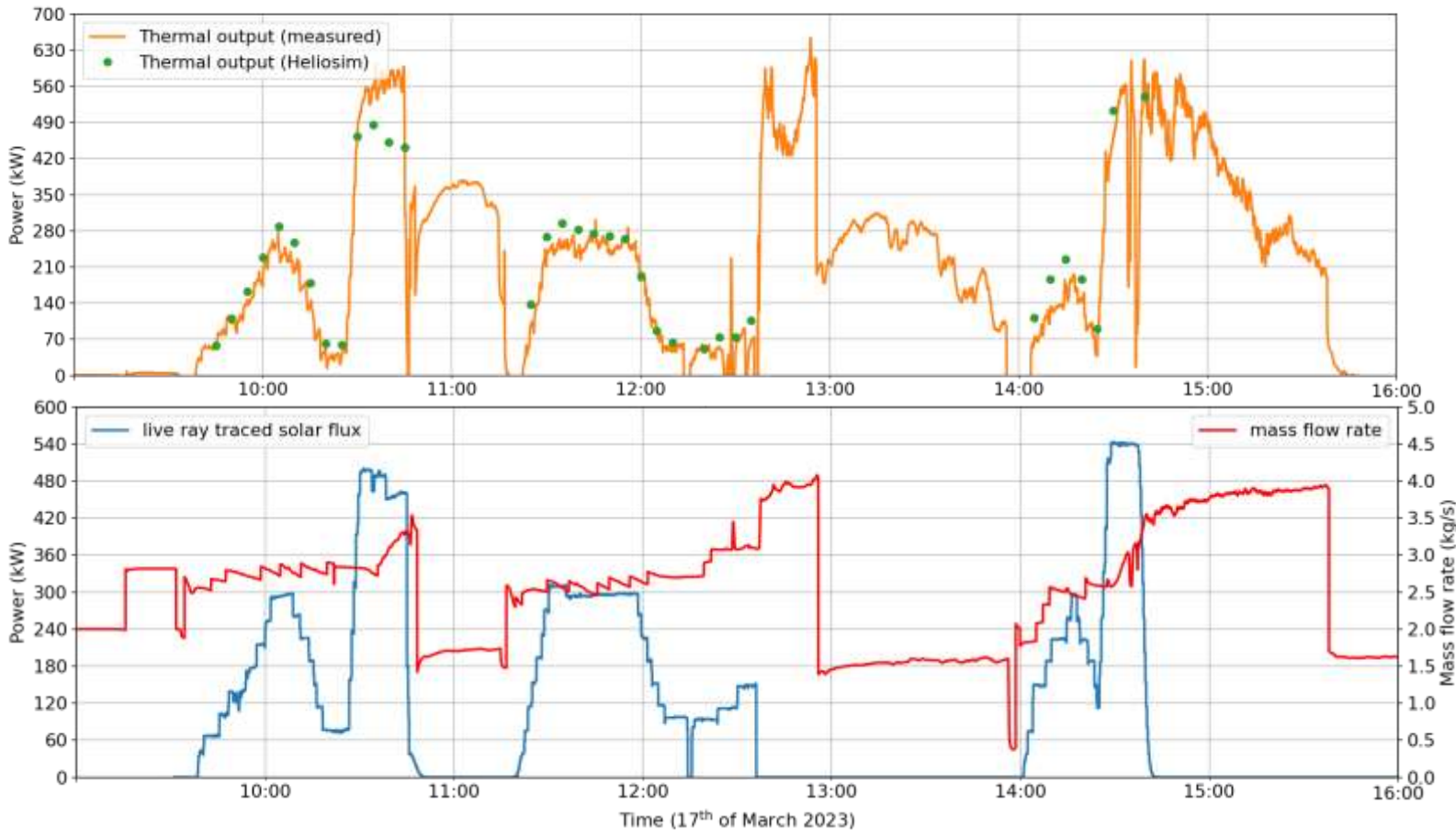


# Results – 17<sup>th</sup> of March 2023



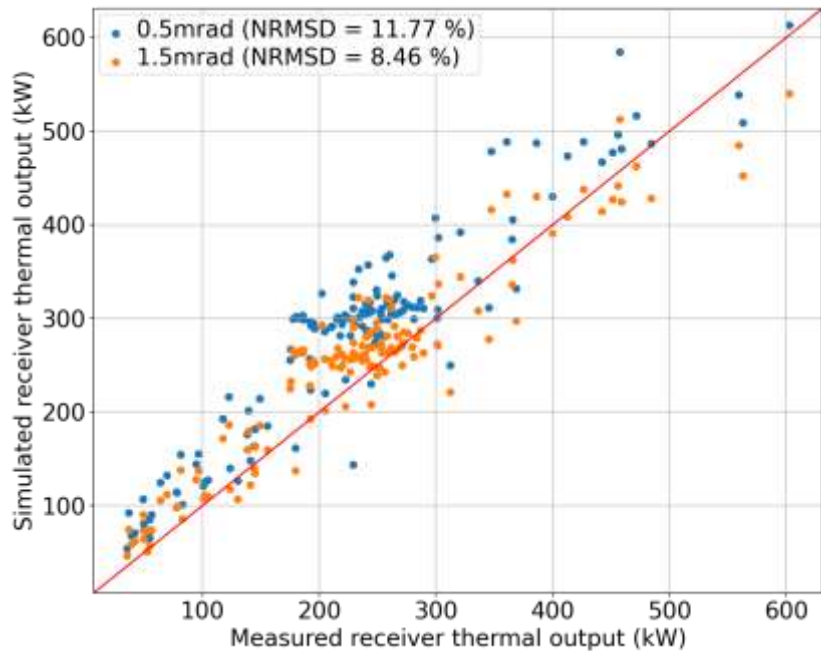


# Results – 17<sup>th</sup> of March 2023

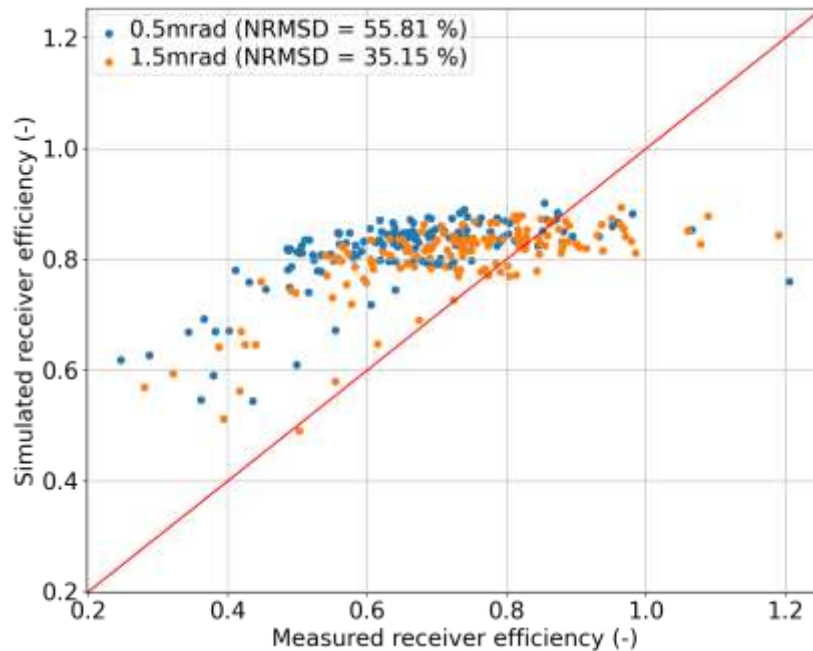




# Aggregated results



(a) Thermal output

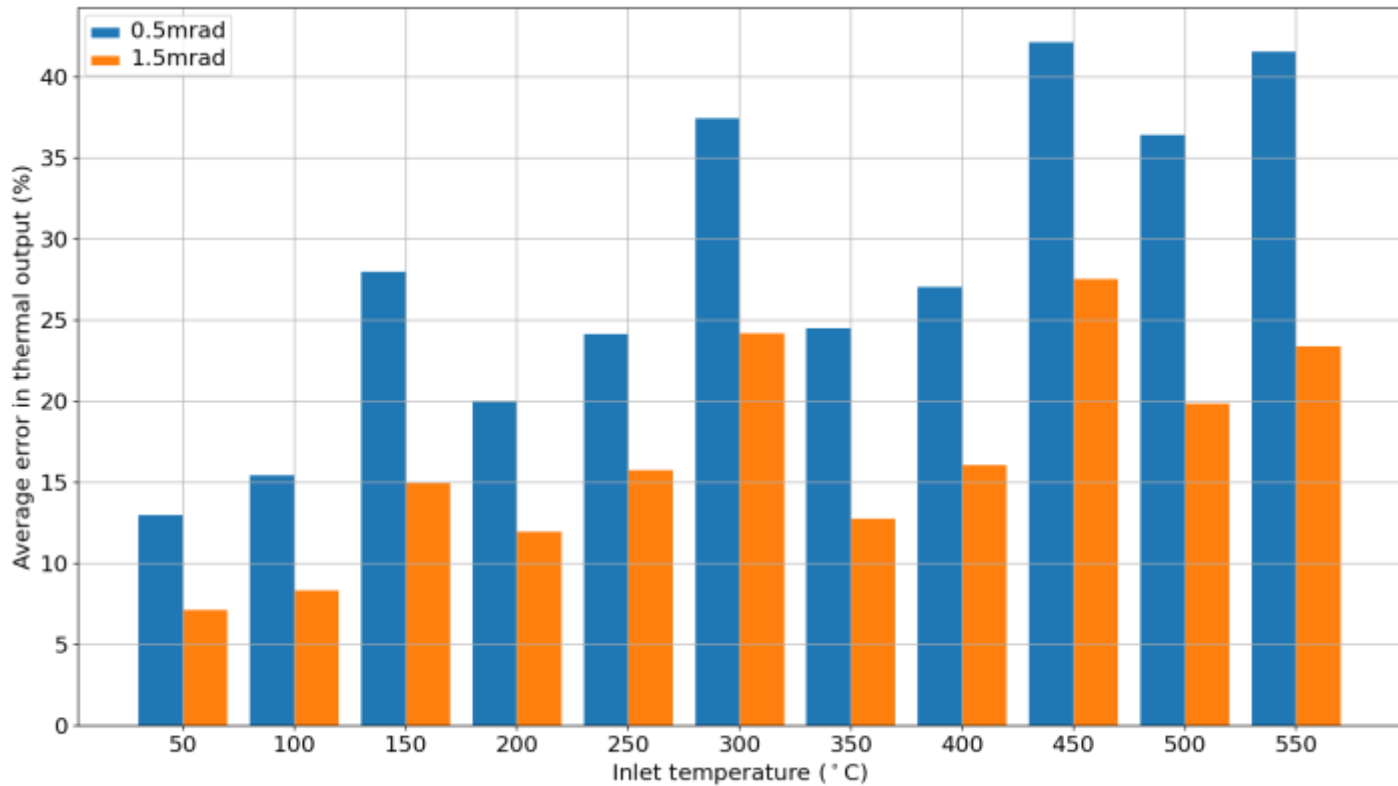


(b) Efficiency

\* “Measured” receiver efficiency is calculated as:  
measured thermal output / simulated aperture solar flux



# Aggregated results





# Conclusion

- A computational model for the ASTRI demonstration particle receiver has been developed using the Heliosim software
- The model has been applied to rebuild single-bin experiments conducted between May 2022 and March 2023
- Reasonable agreement is observed between measurement and simulation for the timeseries trend of receiver outlet temperature
- Significant discrepancy is found for receiver thermal output and receiver efficiency, especially at elevated inlet temperatures (above 400°C)
- Both the measured particle mass flow rate and aperture solar flux need to be more accurately determined to characterise the receiver performance with more confidence





# Thank you

**Energy**

Daniel Potter

Research Scientist

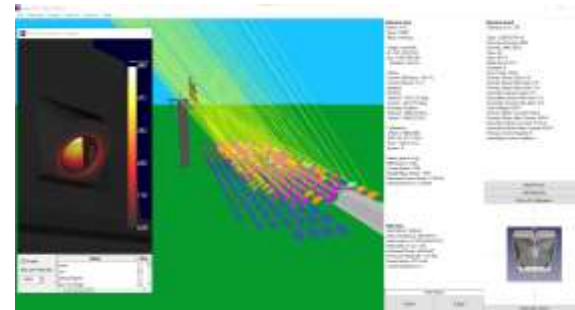
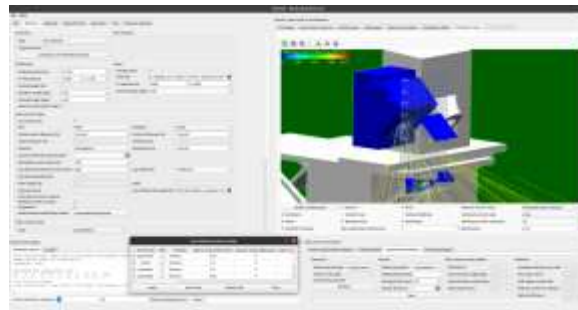
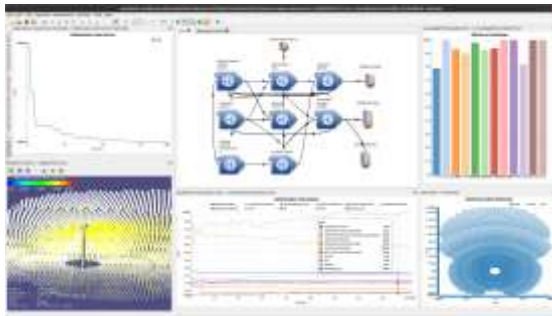
+61 2 4960 6004

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# The Heliosim software family

A collection of libraries, plugins and applications that has been developed by the CSIRO for modelling, design, and control of heliostat-based concentrating solar thermal (CST) systems [1, 2]



## Defining features

- Object-orientated implementation of core functionality in C++
- GPU-accelerated Monte Carlo ray tracing with mesh-based models of heliostats, receivers, etc.
- Standalone applications with graphical and command line interfaces

## Usage

- Provides critical tools for the design and operation of receiver experiments at CSIRO Newcastle
- Modelling capability is used in numerous research and commercial projects
  - Detailed performance modelling of heliostat fields & receivers
  - LCOE & LCOH based system-level optimisation

1. D. Potter, et al., AIP Conf. Proc., vol. 2033, 2018, doi: 10.1063/1.5067213
2. D. Potter, "The Heliosim software family: modelling, design, and control tools for CST systems," in Proceedings of the Asia Pacific Solar Research Conference, 2022.