

A STUDY ON THE TOLERANCE OF PEROVSKITE SOLAR CELLS

UNDER PROTON RADIATIONS

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Perovskite solar cells



- High power conversion efficiency
- Simple and low cost solution-based processes
- High specific power (power per weight)
- High tolerance under high energy particle radiations





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Unstable at ambient condition (oxygen and moisture)

Why space application?



Price optimization for commercialization



• Unstable with oxygen and moisture



- High tolerance under particle radiations
- High PCE
- High power per weight

Transportation cost

High energy particle radiations



Cosmic rays: γ , e⁻, **p**⁺, β ⁺...

Among them, p+ is the main source of damage due to abundance and energy (mass)

Defect creation within PSC's functional layers

→ This work will focus on defect creation and efficiency degradation under proton radiation

Simulation of defect creation within perovskite



- SRIM simulation for defect production profiles
- COMSOL simulation for JV-curves (with additional defects from SRIM)

For detail, please see Nguyen, D.T. et al. (2023), Adv. Energy Sustainability Res. 2300085 ANU School of Engineering | A Study on the Tolerance of Perovskite Solar Cells under Proton Radiations

Simulation of efficiency degradation

J-V changed under 10MeV proton radiation



Experimental setup

- 10MeV proton radiation
- Total irradiation area is 20x20mm which overscans the sample's size of 12.5x14mm
- Flux 1e9 p/cm2.s to avoid heating
- Fluence: 1e12, 1e13, 1e14 p/cm2; in which 1e14 p/cm2 equivalents to ~5.5 years in LEO or ~100 years in GEO ^(*)

(*) Calculated based on 10 MeV proton equivalent dose, source: Brandon K. Durant et al. ACS Energy Letters 2021, 6, 7, 2362-2368



Samples including:

2 cells, 2 films, and one ITO glass for each condition

Reference SEM



1e12 p/cm² SEM







PL



J-V measured after radiated



Simulation vs. Experiment

Conclusion

- → Mechanical damage and ions invasion into next layers after being irradiated
- → Samples remain 89% of their initial efficiency after ~100 years in GEO
- Experiment well matches simulation results
- Further study in space applications: **low light & low temperature (deep space)**



THANK YOU!

Contact Us

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(Paper under preparation)









Supporting slide

Structure	Energy	Fluence	Initial effi.	Remaining	Ref.	
ITO/TiO ₂ /MAPbl ₃ /P3HT/Au	150keV	1e12	4.8%	95%		
		1e13		80%	1	_
		1e14		70%		
ITO/c-TiO ₂ /mp-TiO ₂ /MAPbI _{3-x} Cl _x /P3HT/Au	50keV	1e12	4.8%	100%	2	
		1e13		104%		
		1e14		110%		
ITO/polyTPD/PFN- Br/FA _{0.8} Cs _{0.2} Pbl _{2.4} Br _{0.6} Cl _{0.02} /C60/SnO _x /ZTO/ITO/Al ₂ O ₃	50keV		5.6%	77%	3	
	300keV	1e12	5.8%	147%		[1
	2.5MeV		5.4%	117%		4
ITO/SnO ₂ /Cs _{0.15} FA _{0.85} Pbl ₃ /PTAA:Spiro/Au	7MeV	1e13	16.3%	48%	4	[2
ITO/SnO ₂ /Cs _{0.15} FA _{0.85} PbI ₃ /PTAA/Au			14.9%	57%		v
ITO/SnO ₂ /Cs _{0.15} FA _{0.85} Pbl ₃ /PTAA:C8BTBT/Au			18.0%	57%		[3
ITO/polyTPD/PFN/FA _{0.6} Cs _{0.3} DMA _{0.1} Pbl _{2.4} Br _{0.6} /LiF/C60/PEIE/	68MeV	2e12	15.6%	97%	5	
AZO/ITO/PEDOT:PSS/FA _{0.75} Cs _{0.25} Sn _{0.5} Pb _{0.5} I ₃ /C60/BCP/Au		1e13		94%		[2
ITO/PEDOT:PSS/MAPbl ₃ /PC61BM/BCP/Ag	68MeV	1e12	4.7%	121%	6	E
ITO/SnO ₂ /Cs _{0.1} FA _{0.9} Pbl ₃ /Spiro/Au	10MeV	1e12	24.1%	99%	This	[5
		1e13		98%		2 [4
		1e14		89%		N



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